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# New frontiers of altimetry

Lake Constance - Germany,  
27-31 October 2014

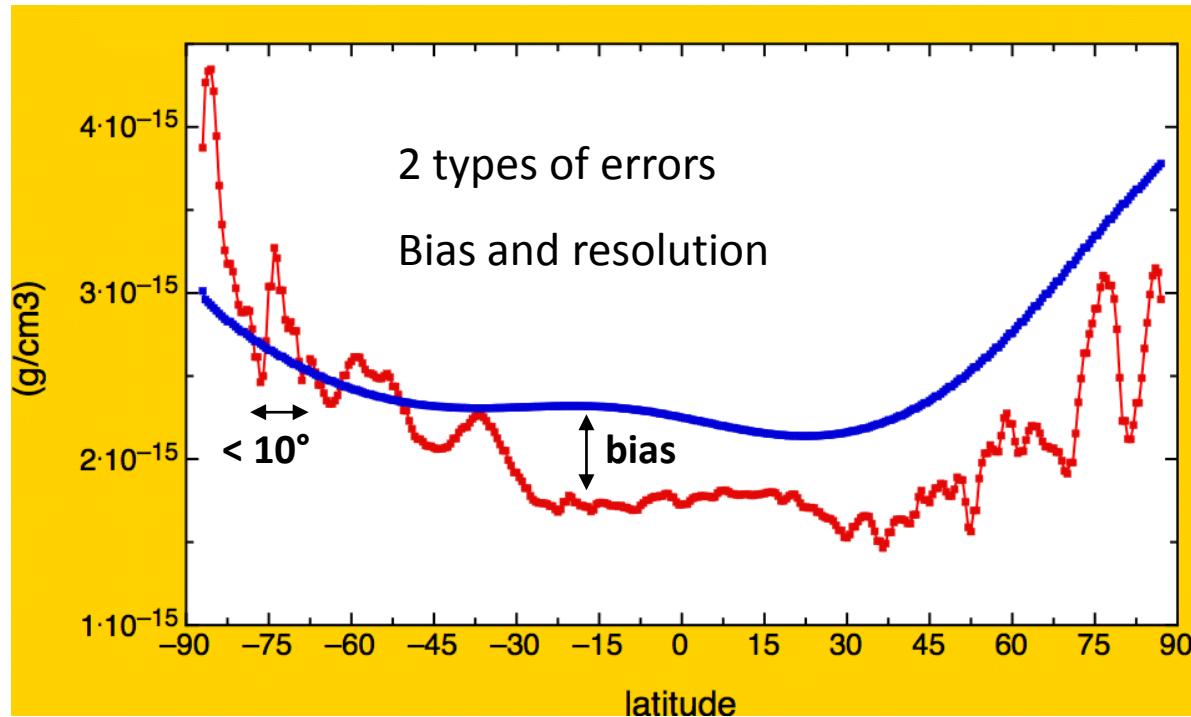
## The semi-empirical thermosphere model DTM2013 (Drag Temperature Model)

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# ATMOP (FP7 project): objective

To improve the semi-empirical DTM model in the 250-1000 km altitude range of the upper atmosphere and reduce systematic errors in particular (most important for orbit computation), at least to the level of JB2008



# DTM model

The following variations are modeled (temperature and composition)

## periodic

diurnal

semi-diurnal

ter-diurnal

annual

semi-annual

(+ coupling with F30)

## non-periodic

latitude (zonal)

**solar activity: F30**

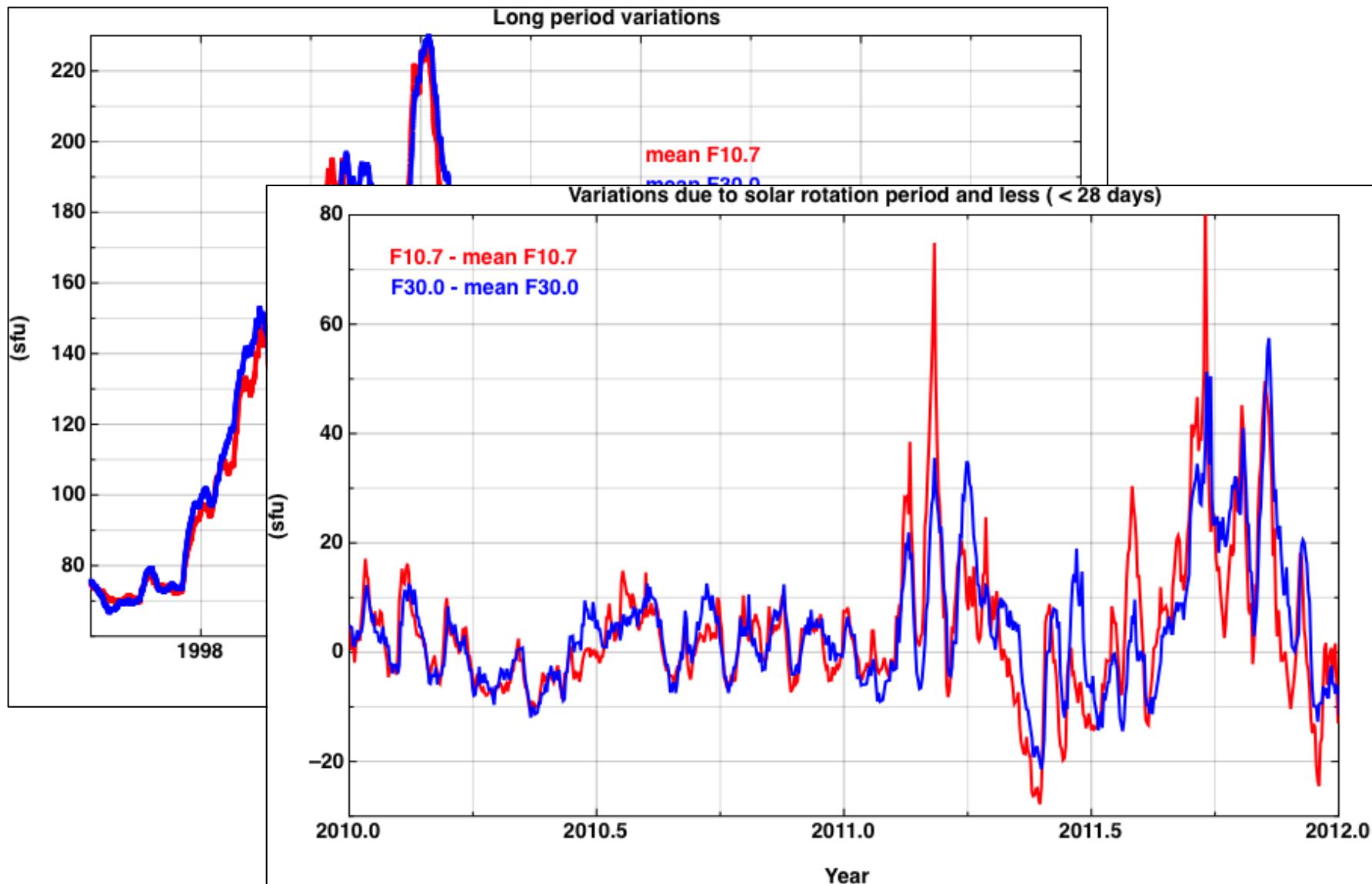
geomagnetic activity: kp

## Pros/Cons of empirical models:

*Fast computation, easy to use, relatively accurate (10-20%, 1 $\sigma$ )*

*Low resolution, static response, proxy use, scale/bias problems,....*

# Solar proxies: $F_{30}$ and $F_{10.7}$



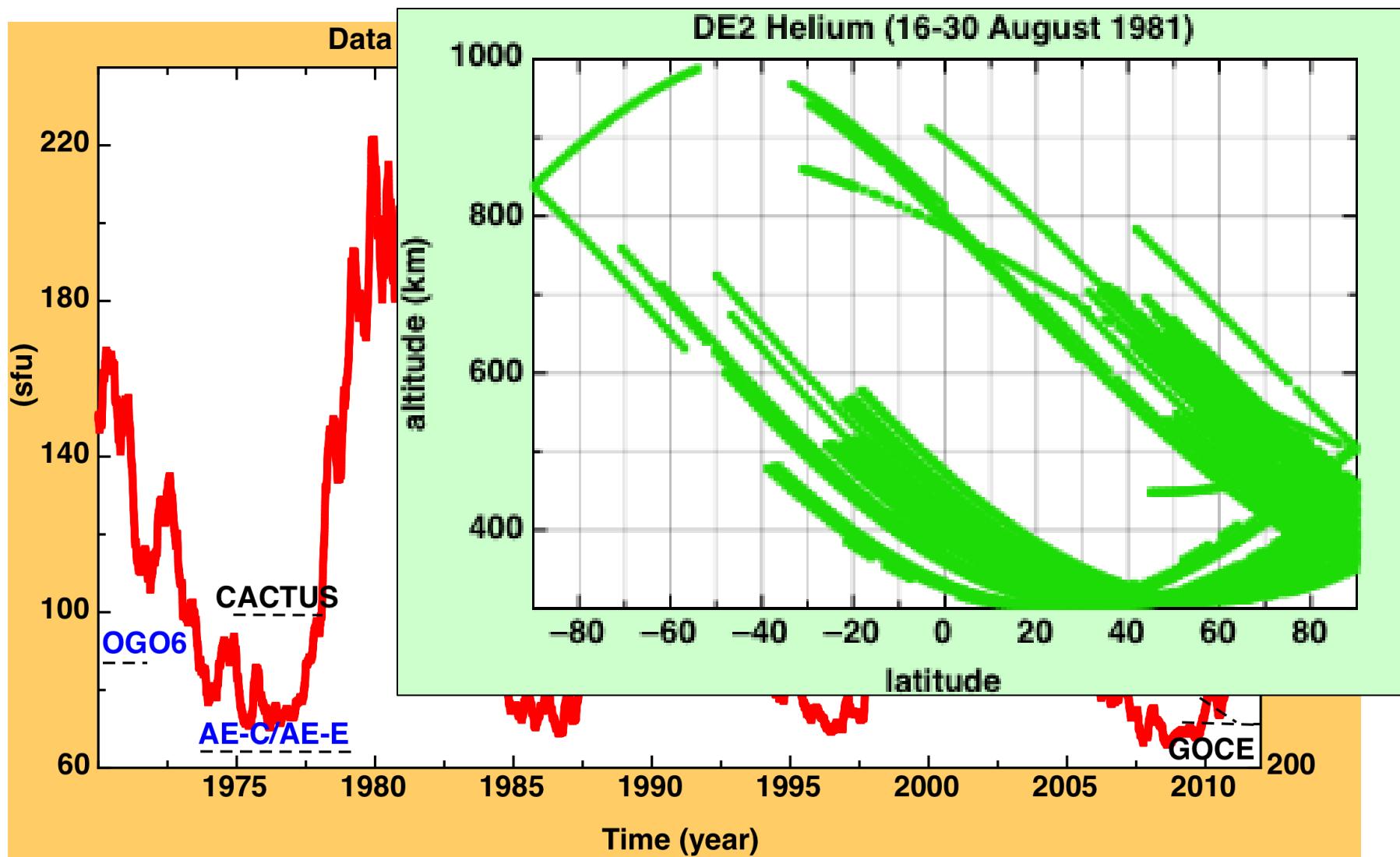
# DTM model

## Data used in the construction of DTM2013:

✓ CHAMP	05/2001 - 08/2010
✓ GRACE	01/2003 - 12/2011
✓ GOCE	11/2009 - 05/2012
✓ Starlette & Stella	01/1994 - 12/2012
✓ Deimos-1	03/2010 - 09/2011
✓ CACTUS	07/1975 - 01/1979
✓ OGO6	06/1969 - 08/1975
✓ DE-2 (T, He, O, N2)	08/1981 - 02/1983
✓ AE-C (N2)	01/1974 - 04/1977
✓ AE-E (T, He, O)	12/1975 - 05/1981

# DTM model

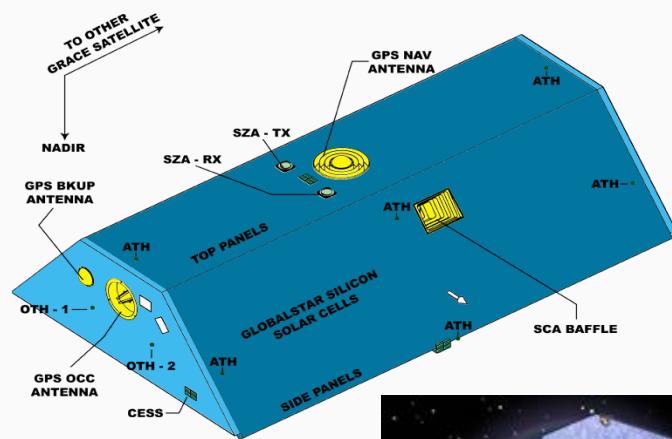
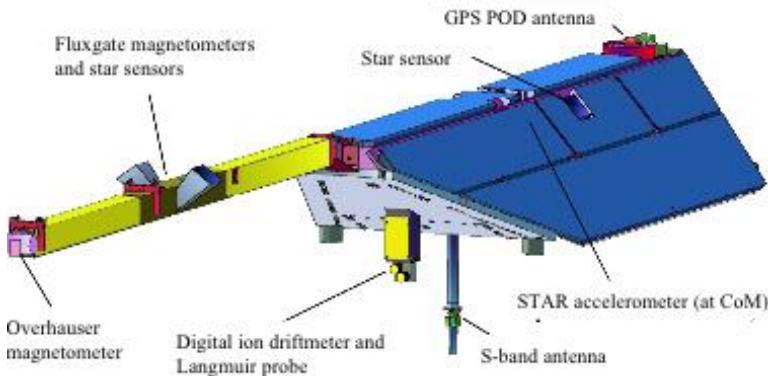
NB: No data above 1000 km; data is sparse at 800 km



# Data: CHAMP, GRACE and GOCE

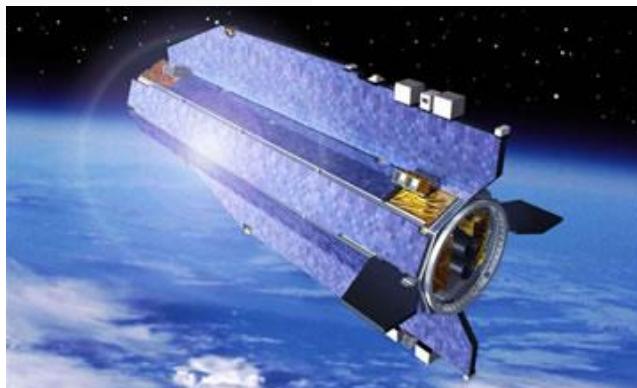
## CHAMP (2000-2010):

- **STAR resolution:**  $3 \cdot 10^{-9} \text{ m/s}^2/\text{Hz}^{0.5}$
- GPS and SLR
- inclination: 87°
- Altitude: 460-300 km



## GRACE (2003-2011):

- **SuperSTAR resolution:**  $1 \cdot 10^{-10} \text{ m/s}^2/\text{Hz}^{0.5}$
- GPS and SLR
- inclination: 90°
- Altitude: 490-450 km



## GOCE (11/2009 – 5/2012):

- **Acc. resolution:**  $1 \cdot 10^{-12} \text{ m/s}^2/\text{Hz}^{0.5}$
- **ion propulsion**
- GPS and SLR
- inclination: 96.5°
- Altitude: 255-225 km

# Calculating drag, deriving density

The drag acceleration:

$$a_{drag} = -\frac{1}{2} C_D \frac{A}{m} \rho v^2$$

$v$  = speed with respect to co-rotating atmosphere

$A$  = surface perpendicular to speed (ram area)

$m$  = mass

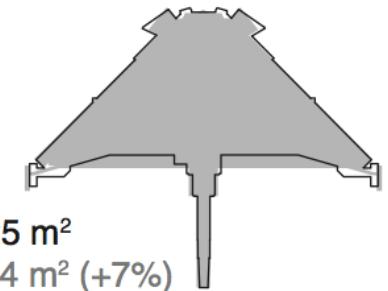
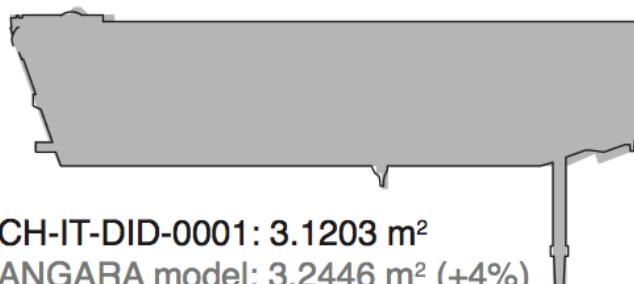
$C_D$  = aerodynamic (or drag) coefficient

$\rho$  = density

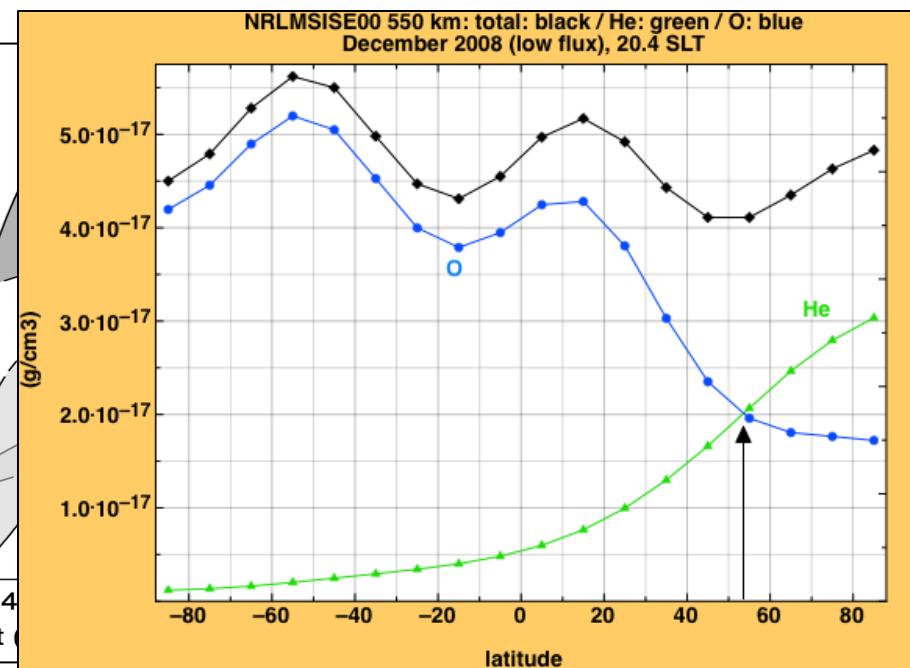
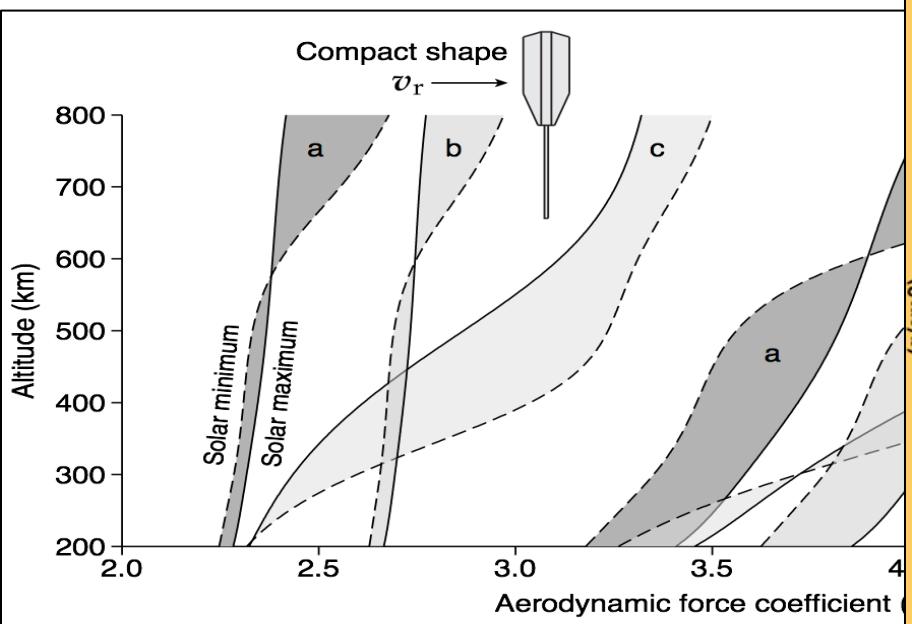
When the drag acceleration is measured/known, and all other parameters are also known, atmospheric density can be computed.

But: errors in  $A$ , and no standard for  $C_D$   « Density? »

# Calculating drag, deriving density



## Problem: no standard for drag computation

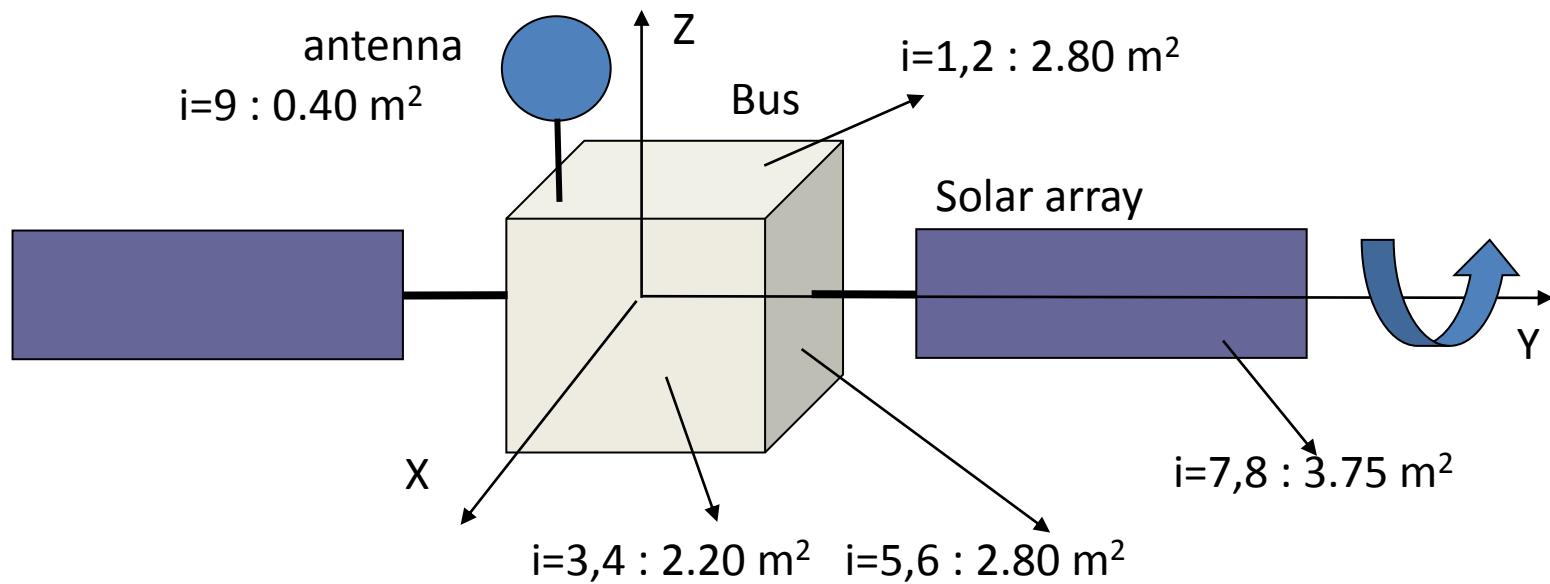


- Drag coefficient (different models and hypotheses)

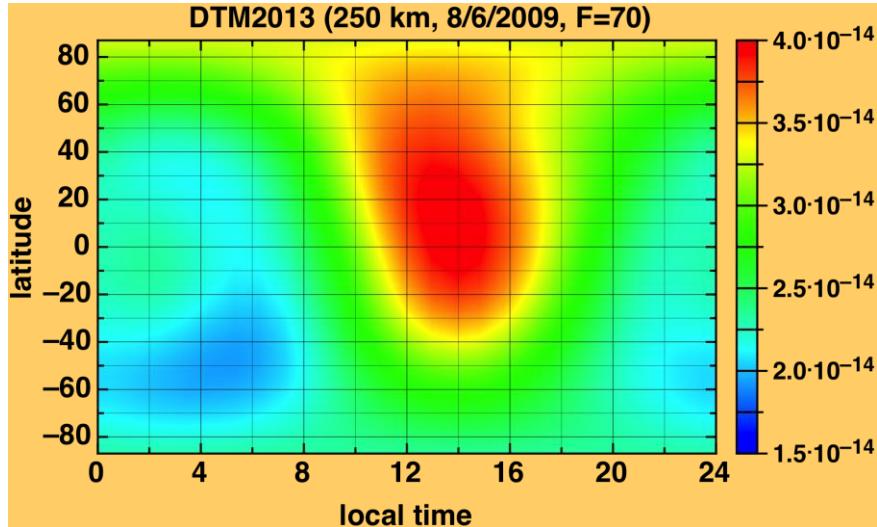
# Calculating drag, deriving density

For EO satellites, the situation is much worse than for CHAMP

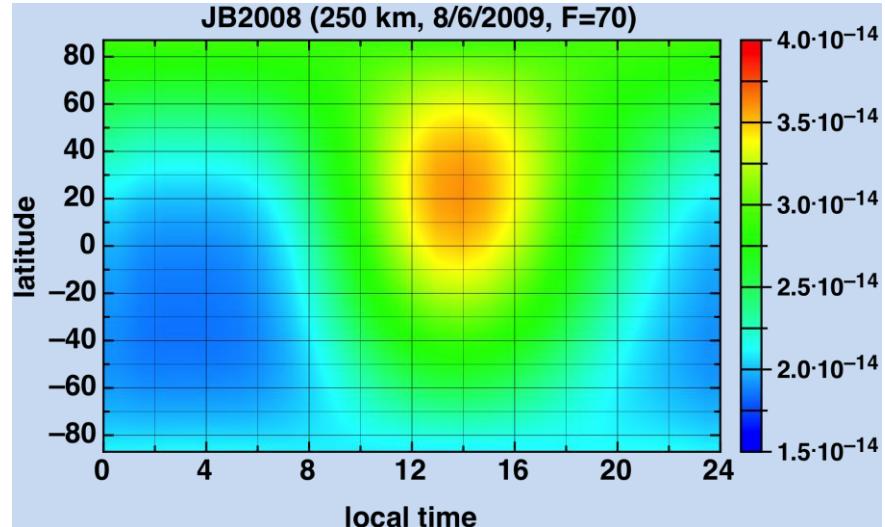
*Simple satellite model (macro model): box and wing(s), antenna*



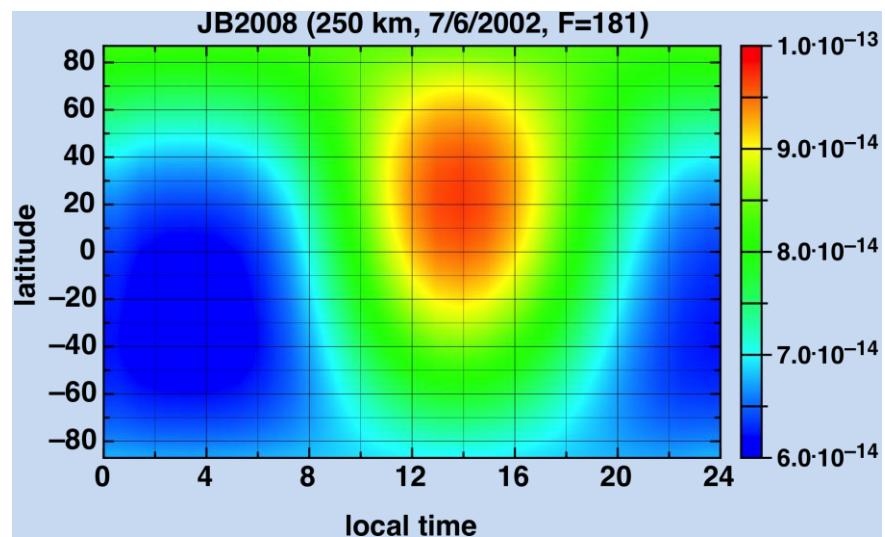
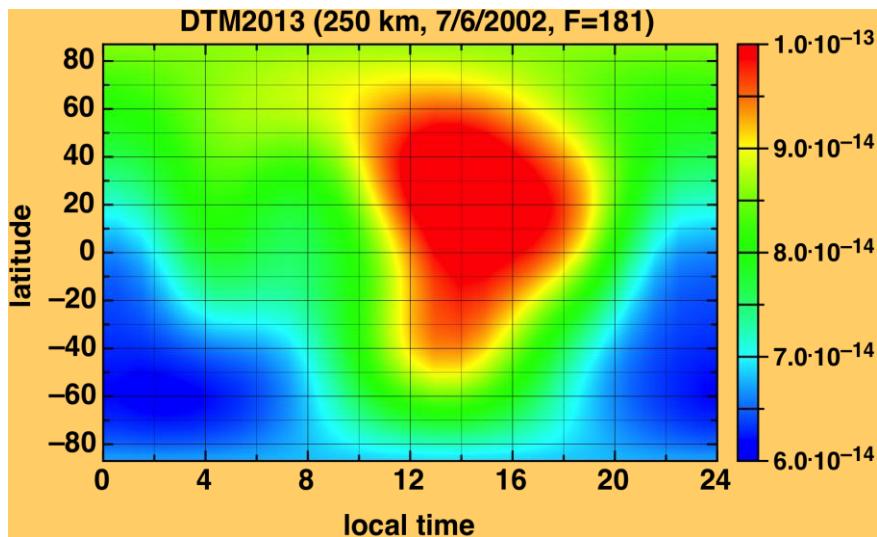
# Model evaluation



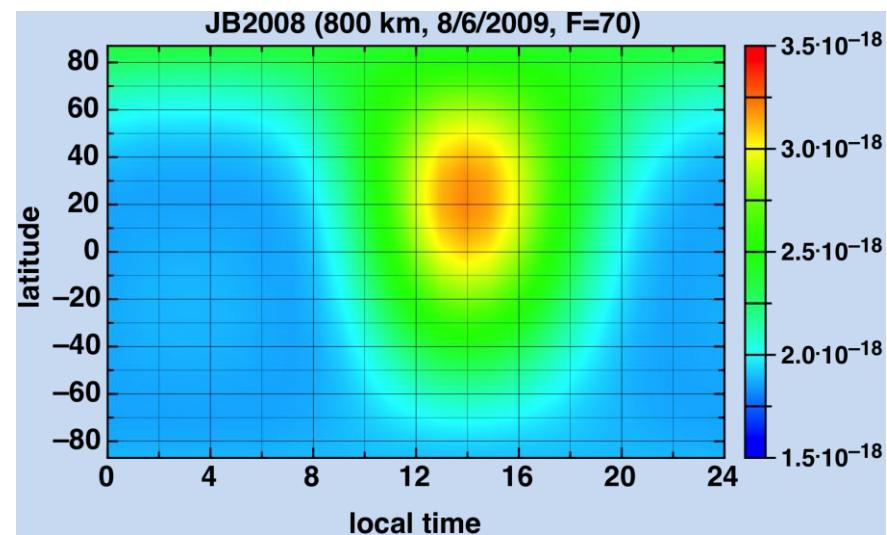
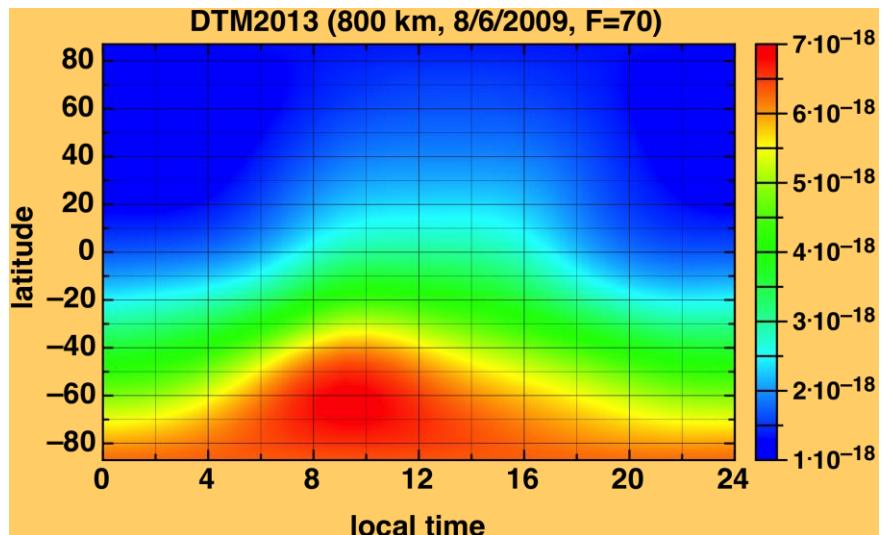
DTM: CIRA optional model



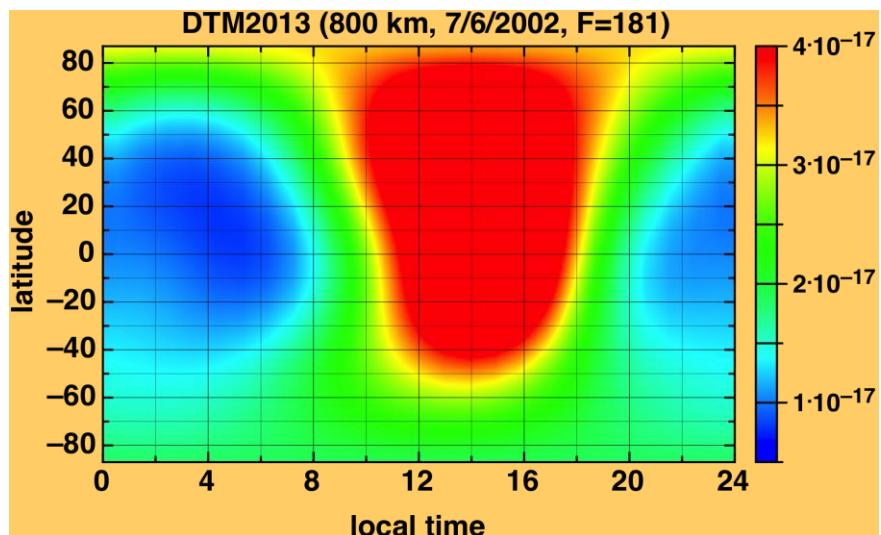
JB2008: CIRA model



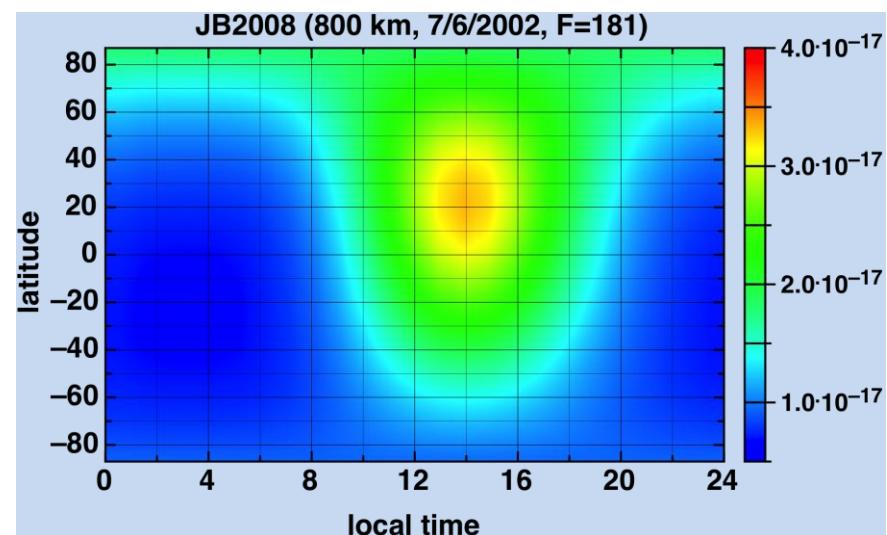
# Model evaluation



DTM: CIRA optional model



JB2008: CIRA model



# Model evaluation

O/C = Observed-to-Modeled ratio (ideally: O/C = 1.00)

Mean of density ratios per year: DTM2009

CHAMP

GRACE

Stella

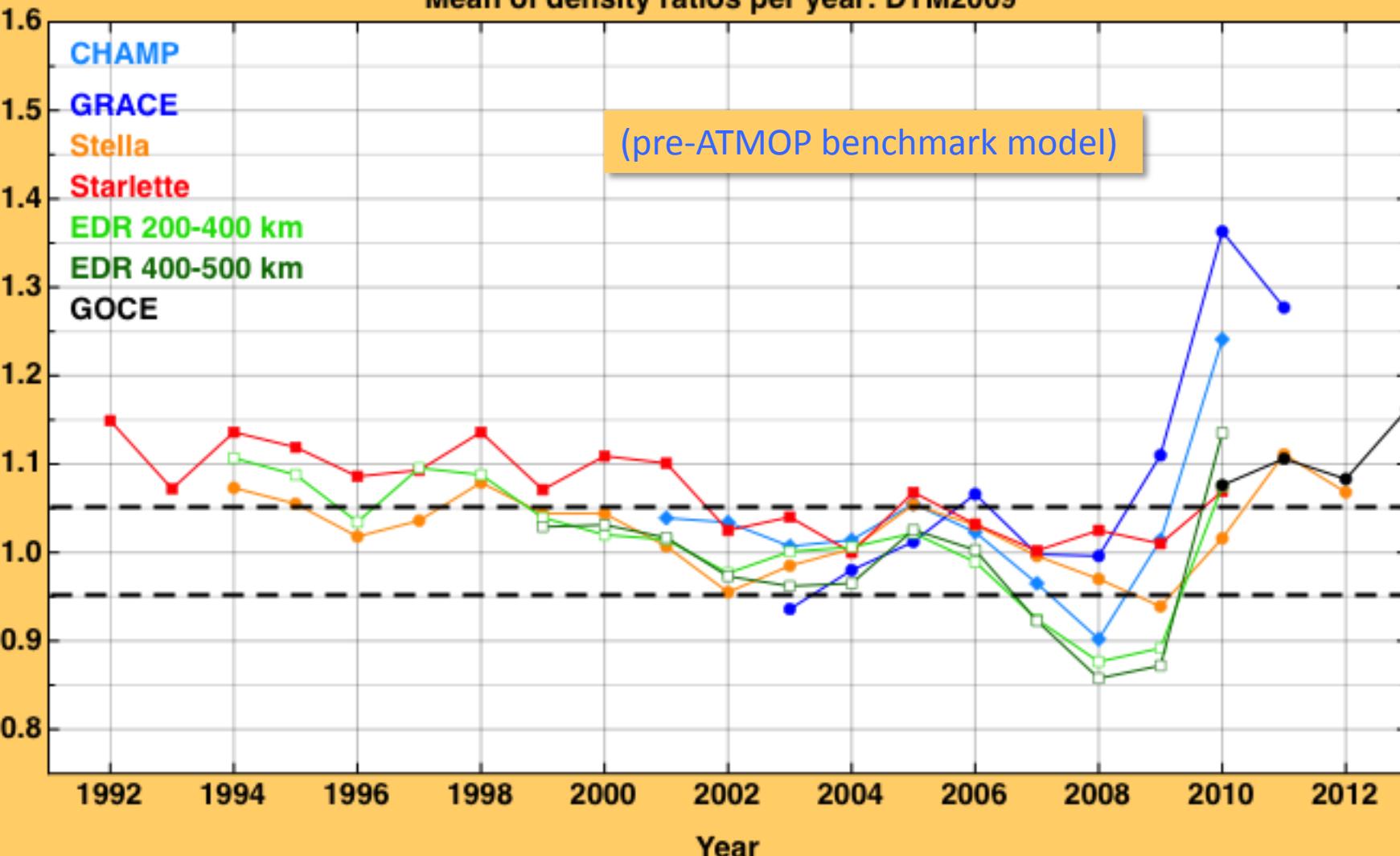
Starlette

EDR 200-400 km

EDR 400-500 km

GOCE

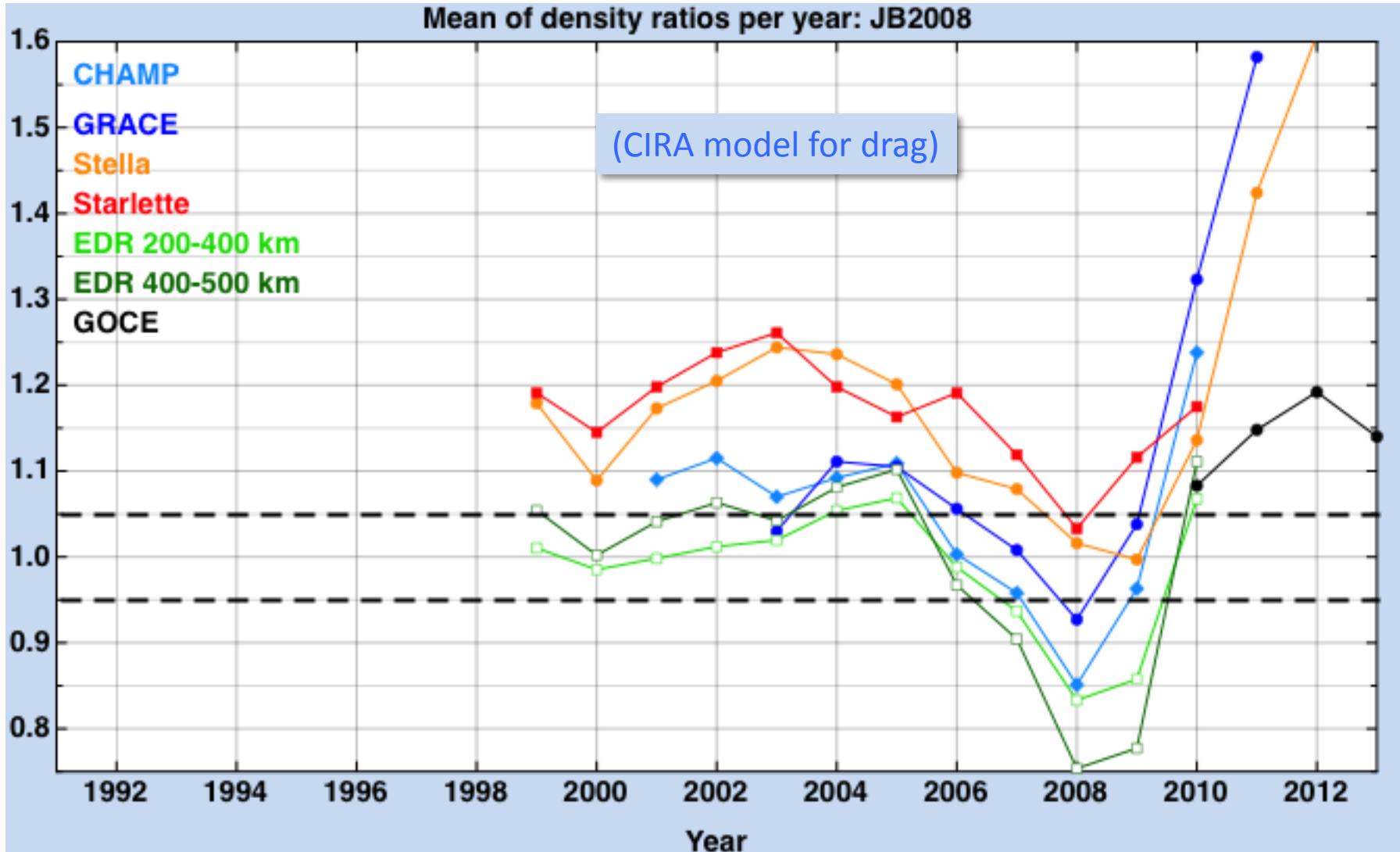
(pre-ATMOP benchmark model)



# Model evaluation

O/C = Observed-to-Modeled ratio (ideally: O/C = 1.00)

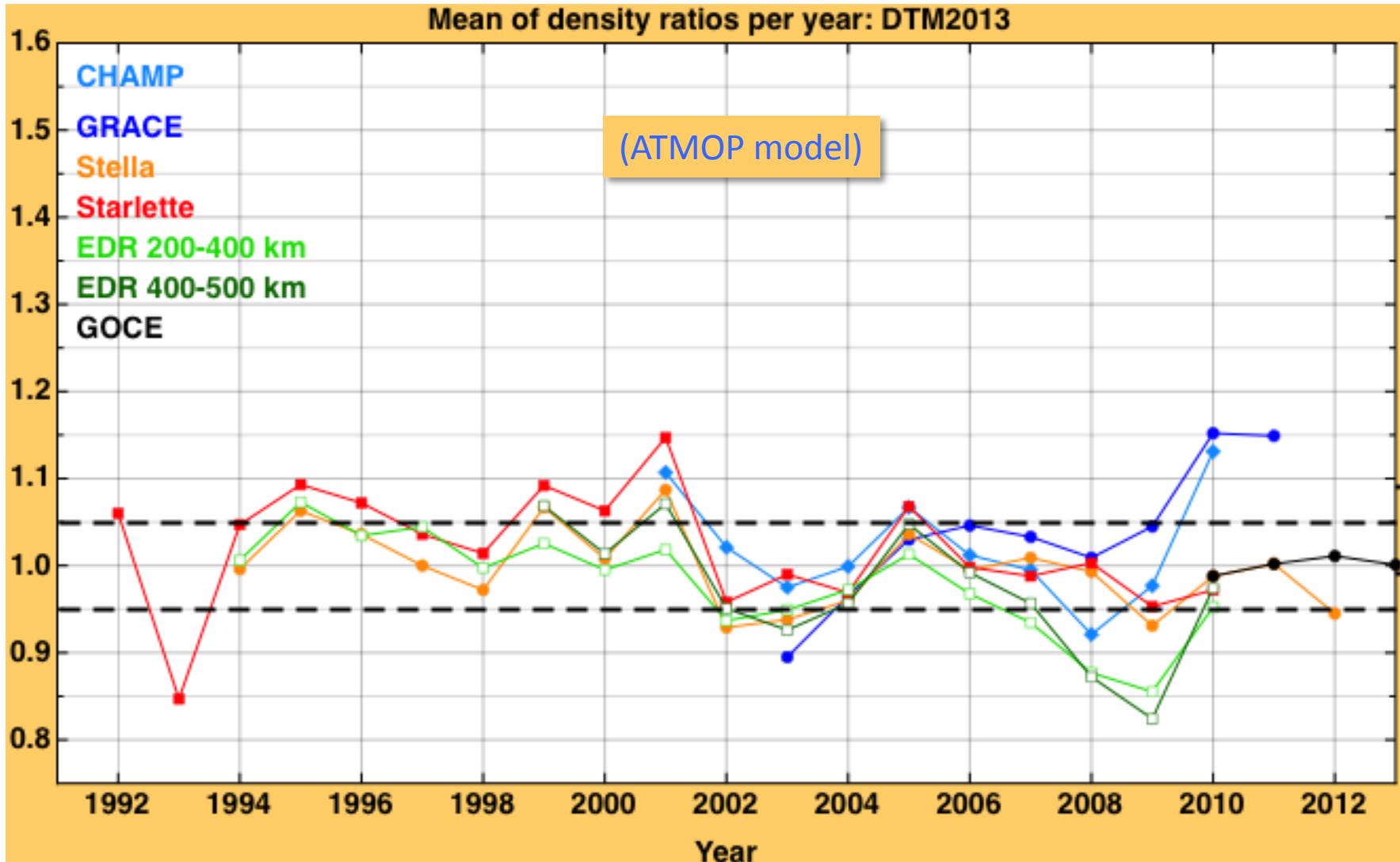
Mean of density ratios per year: JB2008



# Model evaluation

O/C = Observed-to-Modeled ratio (ideally: O/C = 1.00)

Mean of density ratios per year: DTM2013



# Summary and outlook

- ✓ DTM2013 constructed with CHAMP, GRACE, GOCE and historical data
- ✓ F30 solar proxy is used; better results than with F10.7
- ✓ Most precise and least-biased model for GOCE, CHAMP, GRACE, Stella and Starlette
- ✓ Fortran90 code

- *No standard for drag calculation* → *model bias & error*
- *No data above 1000 km* → *pb. TOPEX, Jason*
- *Sparse data above 500 km* → *pb. ERS, Sentinel,..*  
*(not easy to derive from POD)*

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