

---

# DISCUSSION FOR THE MEAN POLE IDS 2017

F. Mercier and A. Couhert

# IERS STANDARDS, POLE AND MEAN POLE

---

Objective :

response of the earth to the accelerations due to pole changes

Stations positions :

low frequency : postglacial rebound part (linear motion) included in the ITRF  
still correct hypothesis ?

response to annual/chandler frequency band, how to compute it correctly ?

Earth gravity field : effects on J2 orientation (C21,S21)

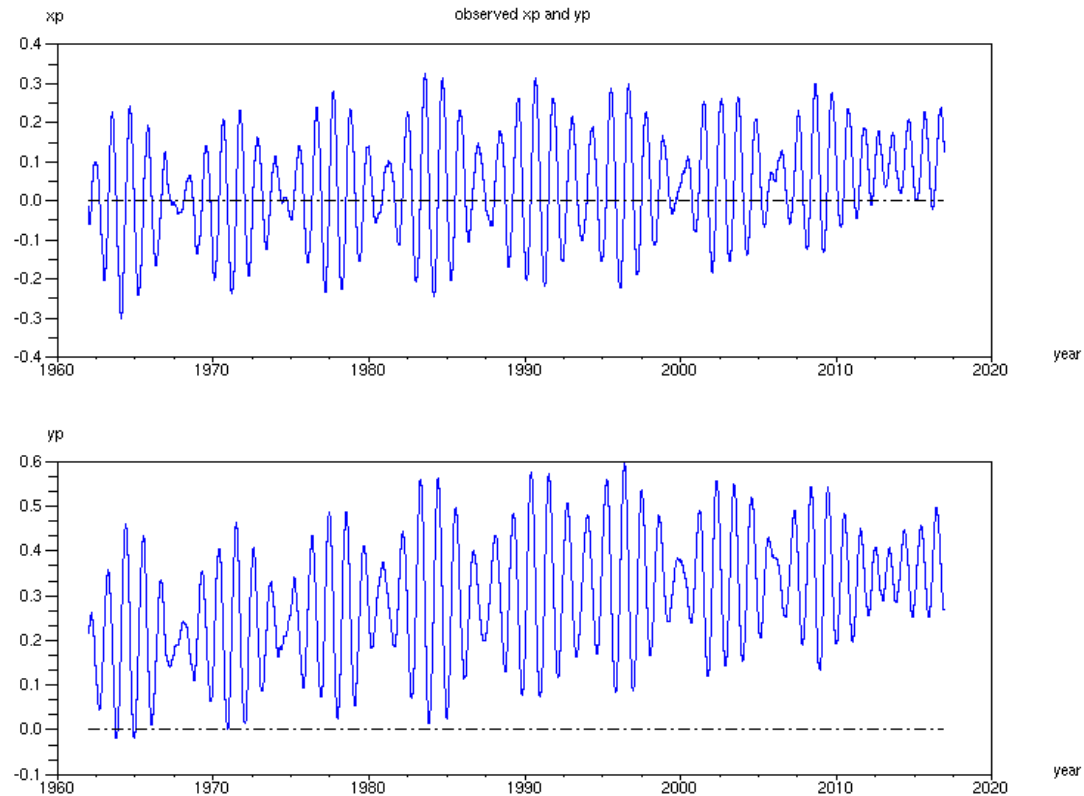
low frequency part : 'mean pole' geometric transformation

annual/Chandler frequency band

# OBSERVED POLAR MOTION

## Main frequency content

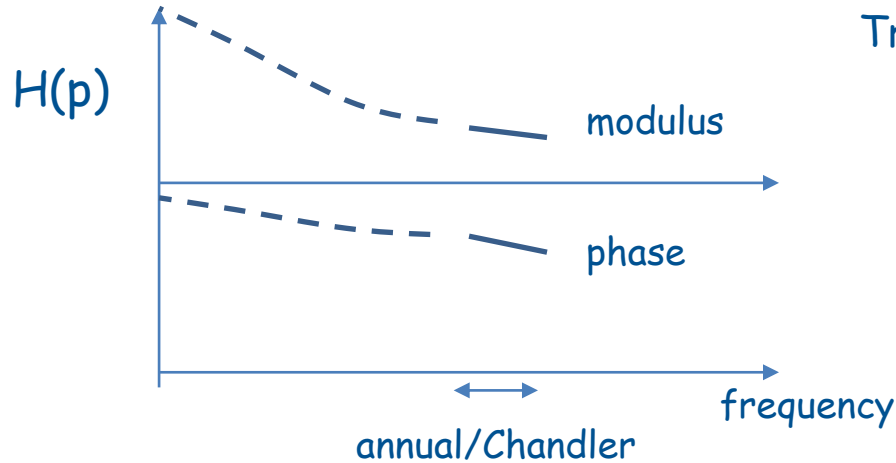
- annual ( ~ 365.25 days)
- Chandler (~ 435 days )
- low frequency  
assumed linear motion up to ~1990  
pluriannual content observed on the  
complete interval



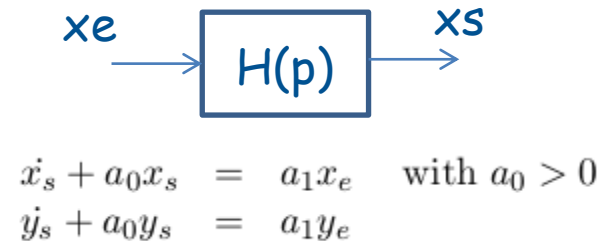
How to estimate the earth response to the complete excitation  
defined as a function of time ?

Remark : the low frequency part of the response may depend on past effects, not known

# DYNAMIC SYSTEM TRANSFER FUNCTION



Transfer function known in the annual/Chandler frequency band



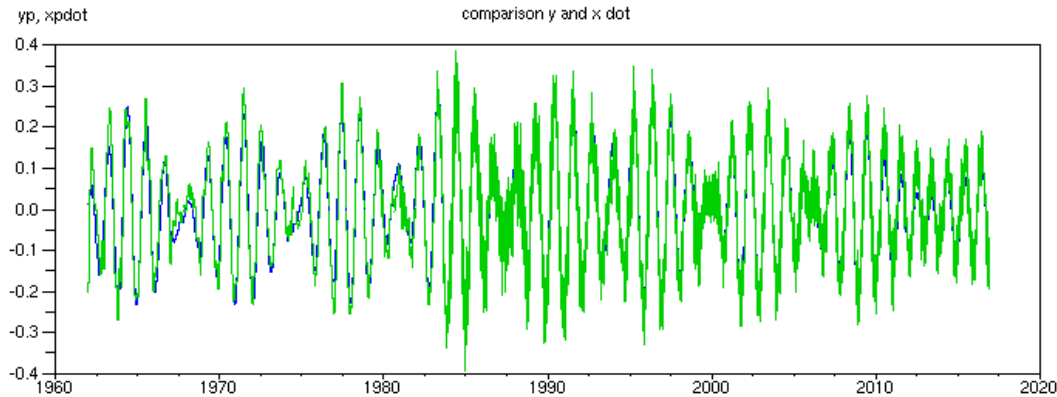
Annual frequency band only :

$$x_e = a \cos(\omega t + \phi) \longrightarrow \begin{cases} (i\omega + a_0)x_s = a_1 x_e \\ (i\omega + a_0)y_s = a_1 y_e \end{cases}$$

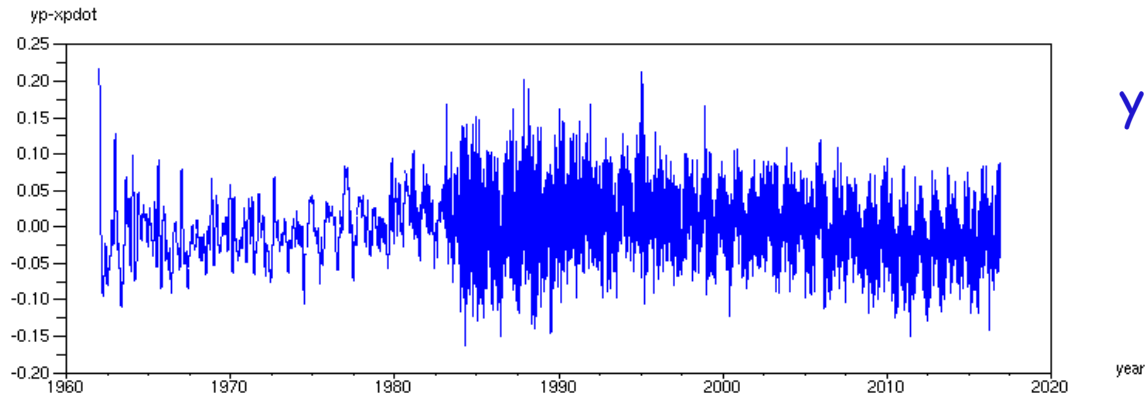
Transformed, for a prograde (clockwise) motion for  $x_e, y_e$  ( $\begin{matrix} \dot{x}_e \leftrightarrow \omega y_e \\ \dot{y}_e \leftrightarrow -\omega x_e \end{matrix}$ ):

$$\begin{cases} x_s = x_e - 0.0115y_e \\ y_s = y_e + 0.0115x_e \end{cases} \quad \text{The formula is valid only in the annual frequency band and for prograde } x_e, y_e$$

# XPDOT AND YP



yp  
xpdot

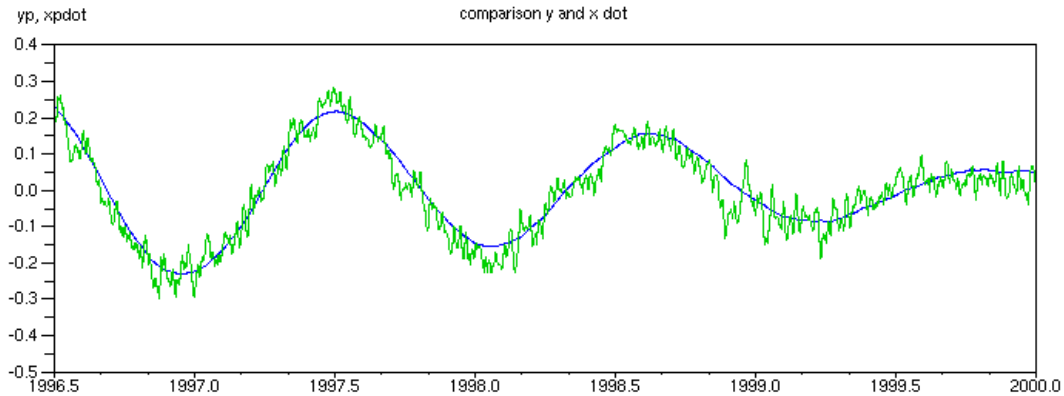


yp - xpdot

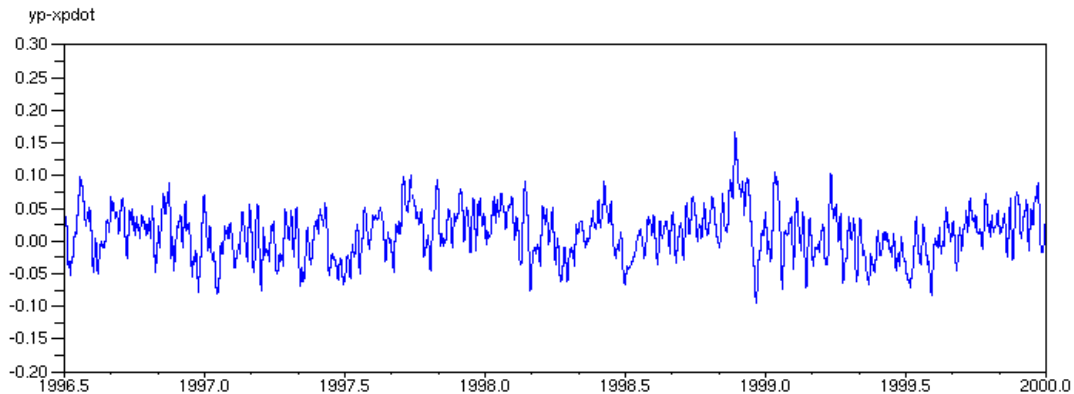
High frequency content due to xpdot estimation method (finite differences)

Low frequency content is from yp (comparison with the complete yp signal)

# XPDOT AND YP



yp  
xpdot



yp - xpdot

High frequency content due to xpdot estimation (finite differences)

xpdot can be efficiently estimated using yp information at annual frequency  
(same property for ypdot and -xp)

# IERS STANDARDS FORMULAS

The formula can be used only in the annual/Chandler frequency band :

in this case, using functions  $\hat{x}_e, \hat{y}_e$  containing all the low frequency content of each signal  $x_e, y_e$  we have :

$$\begin{aligned}x_s &= (x_e - \hat{x}_e) - 0.0115(y_e - \hat{y}_e) \\y_s &= (y_e - \hat{y}_e) + 0.0115(x_e - \hat{x}_e)\end{aligned}$$

$\hat{x}_e, \hat{y}_e$  contain all the contribution outside the frequency band of interest



in the external potential, which is equivalent to changes in the geopotential coefficients  $C_{21}$  and  $S_{21}$ . Using for  $k_2$  the value  $0.3077 + 0.0036i$  appropriate to the polar tide yields

$$\begin{aligned}\Delta \bar{C}_{21} &= -1.333 \times 10^{-9}(m_1 + 0.0115m_2), \\ \Delta \bar{S}_{21} &= -1.333 \times 10^{-9}(m_2 - 0.0115m_1),\end{aligned}$$

where  $m_1$  and  $m_2$  are in seconds of arc.

(IERS standards 2010)

**$m_1$  and  $m_2$  should contain only information in the annual/Chandler frequency band**

$$(m_1 = x_p - \bar{x}_p, \quad m_2 = -(y_p - \bar{y}_p))$$

# IERS STANDARDS FORMULAS

Same remark holds for the stations coordinates :

$$S_r = -33 \sin 2\theta (m_1 \cos \lambda + m_2 \sin \lambda) \text{ inmm,}$$

$$S_\theta = -9 \cos 2\theta (m_1 \cos \lambda + m_2 \sin \lambda) \text{ inmm,}$$

$$S_\lambda = 9 \cos \theta (m_1 \sin \lambda - m_2 \cos \lambda) \text{ inmm,}$$

(IERS standards 2010)

The 'mean pole' to be used in  $m_1$  and  $m_2$  in these formulas is a filtering correction not a geophysical model mean pole

Remark : from Wahr 2015, comparison with standards 2010

$$k_2^r = 0.3077$$

$$k_2^i = 0.0036$$

$$a_e = 6378136.6 \text{ m}$$

$$g_e = 9.7803278 \text{ ms}^{-2}$$

$$\Omega = 2\pi \cdot 1.002737811911/86400$$

$$\begin{aligned} C_{21} &= -\frac{a_e \Omega^2}{g_e \sqrt{15}} (k_2^r m_1 + k_2^i m_2) \\ &= -1.336 \cdot 10^{-9} (m_1 + 0.0117 m_2) \end{aligned}$$

$$\begin{aligned} S_{21} &= -\frac{a_e \Omega^2}{g_e \sqrt{15}} (k_2^r m_2 - k_2^i m_1) \\ &= -1.336 \cdot 10^{-9} (m_2 - 0.0117 m_1) \end{aligned}$$

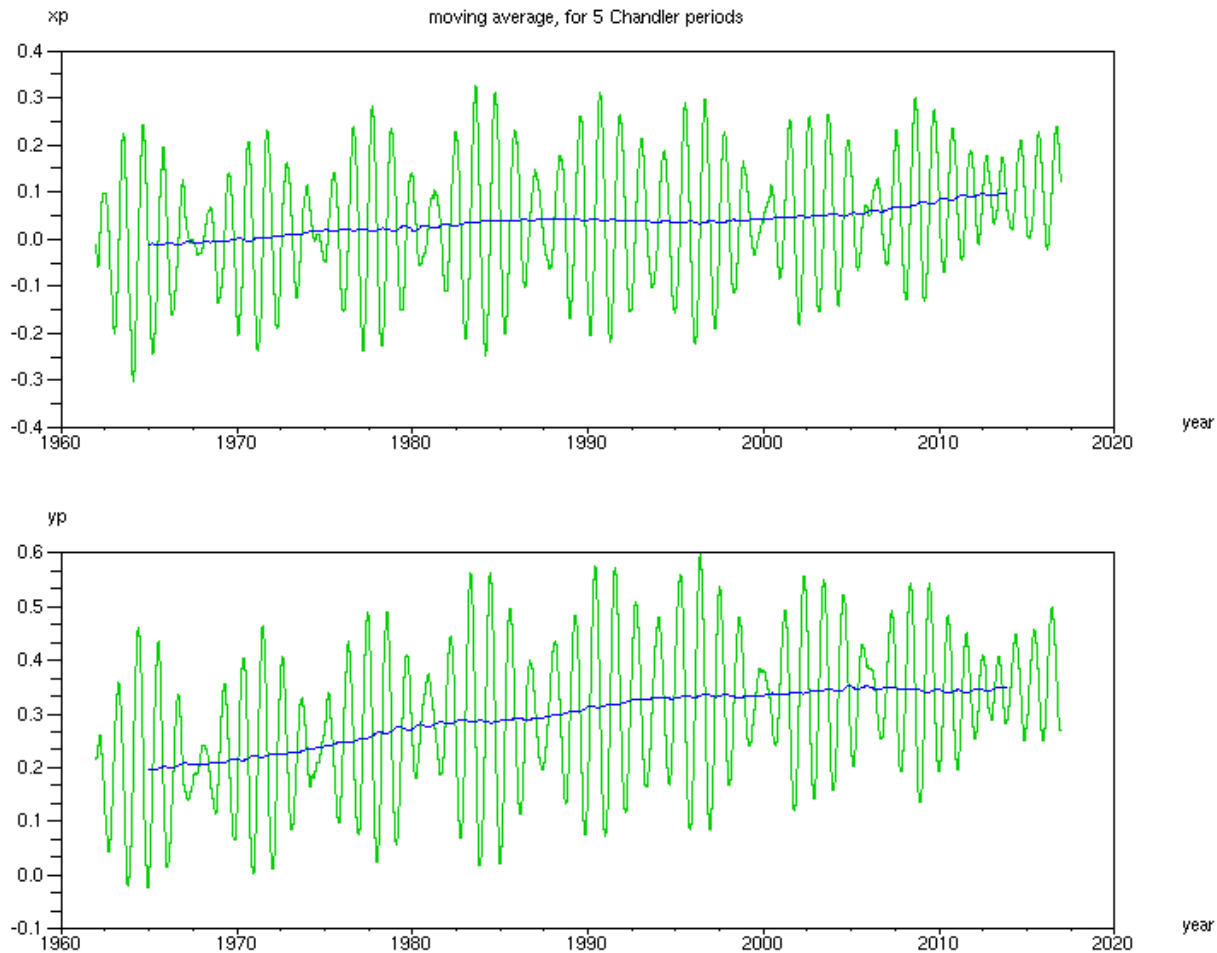
Small differences

0.0117 and 0.0115 (standards)

1.336 and 1.333 (standards)



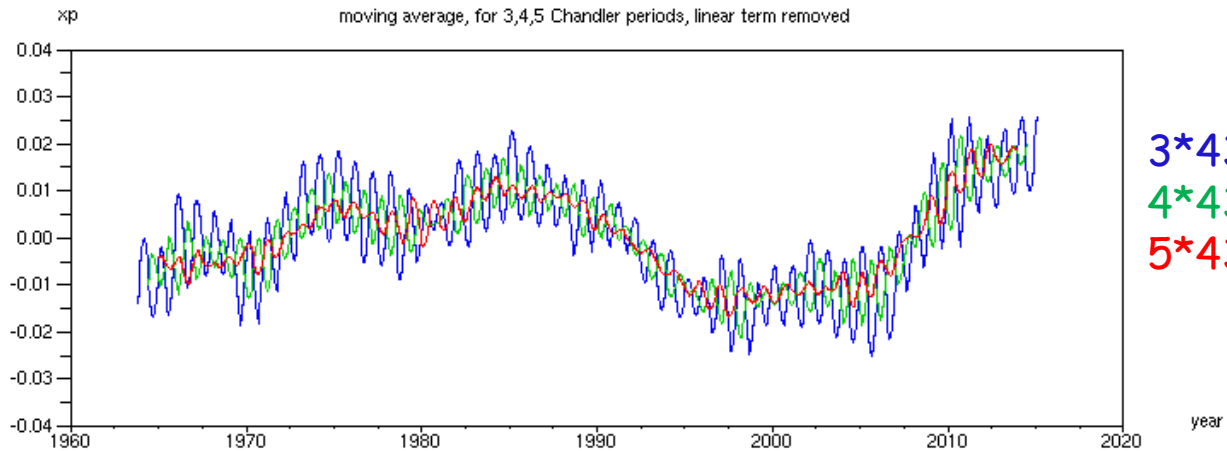
# SIGNAL CHARACTERISTICS OUTSIDE THE ANNUAL BAND



Moving average, 5 Chandler periods : good extraction of the frequency band  
Some pluriannual variations

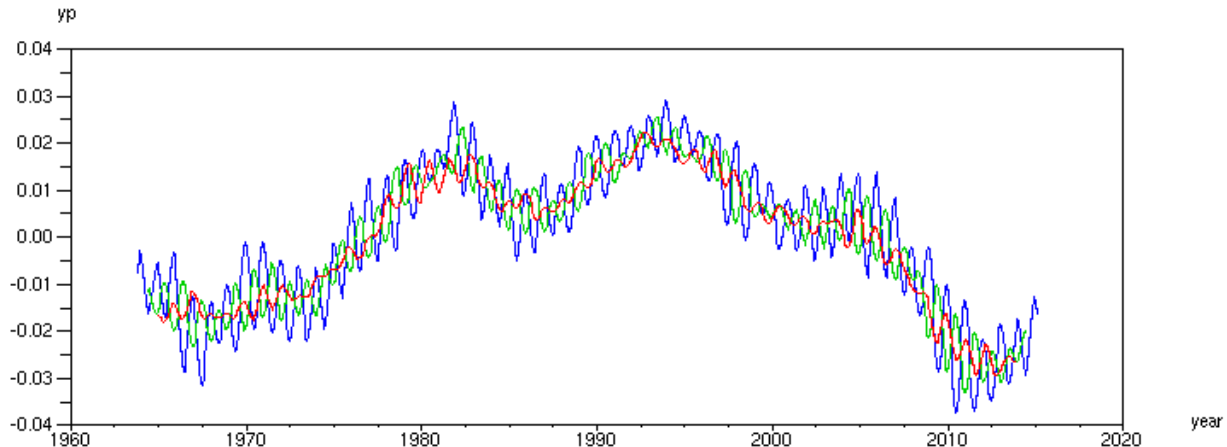
# PRECISION OF THE ESTIMATED LOW FREQUENCY SIGNAL

1 mm effect  
on station positioning  
at annual/Chandler



3\*435 days  
4\*435 days  
5\*435 days

1 mm effect  
on station positioning  
at annual/Chandler



All filtering methods are convenient for a millimetric performance  
The remaining frequencies participating below the annual/Chandler signals have negligible effects  
Using a linear function as filtering reference will produce several millimeters errors

# CONCLUSION

## Annual/Chandler perturbations

m1 and m2 for annual/Chandler formulas in the standards must be computed using a filtered mean pole value for full precision

Filtering with a moving average is sufficient for station positioning

- submillimetric precision for station positioning at annual/Chandler
- linear reference can produce several millimeters errors
- at least, the current standards approach must be used

## Low frequency perturbations

there are pluriannual terms with 0.02, 0.04 arcsec variations

- the response of the earth system at these frequencies is not detailed (use of a 'static' transformation, as in the standards ?)
- are such variations observable in the stations coordinates time series ?
- for the earth potential, the use of a variable gravity field removes the problem (but consistent conventions for the mean pole shall be used)