

Evolution of dynamical orbit model

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Traditional Bernese orbit modeling is based on the empirical and pseudo-stochastic modeling. Dynamical model has been developed in GOP and TUM

Modeling	Empirical-Stochastic (reduced-dynamical)	Dynamical
Satellite attitude and geometry	Not considered	Nominal Box-Wing model
Atmosphere density model	Not applied	MSIS-86
Atmosphere drag	Absorbed by along track stochastic parameters and Y-constant empirical parameter	Scaling coefficient estimated
Solar radiation Pressure	Absorbed by empirical constant parameter in sun-satellite direction	Scaling coefficient estimated or fixed value closed to "1".
Earth radiation	Not applied	A priori model, reflexivity and emissivity
1-per revolution empirical modeling	Sun-Satellite and Y-direction	Along and cross track (optional)

TESTING

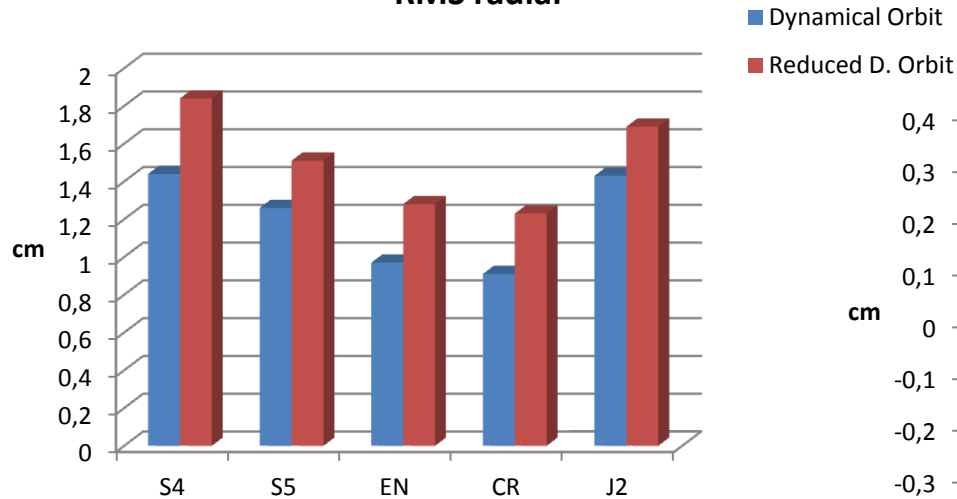
- **First results presented at AGU Fall meeting (december 2011) , detected a signal of 14-15 days**
- **Problem related to ocean tidal modeling, recently corrected**
- **Preliminary comparison with CNES/SSALTO orbits (40 days)**
- **Preliminary comparison of weekly free network solutions (quarter a year)**
- **To be confirmed by longer time series and other tests**

Comparison of the estimated DORIS orbits and CNES/SSALTO multitechnique

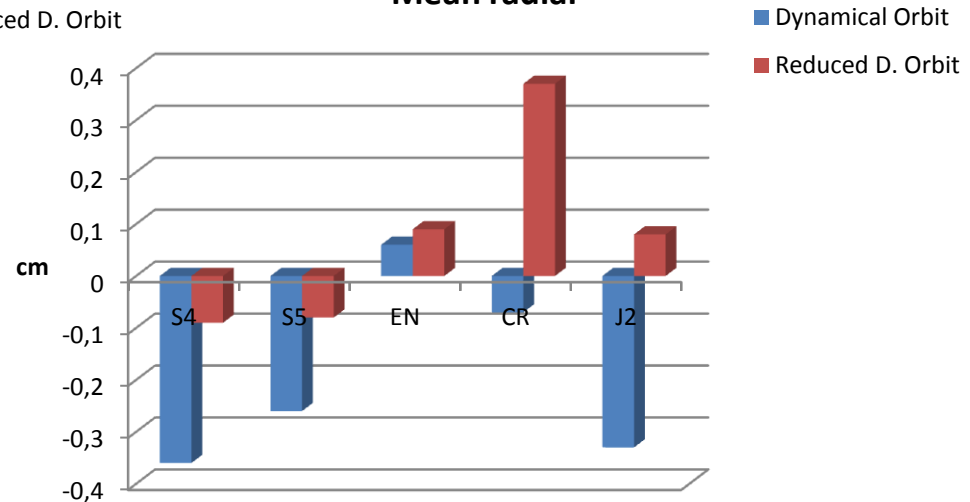
orbits - RADIAL direction

Days 001-040 of 2011

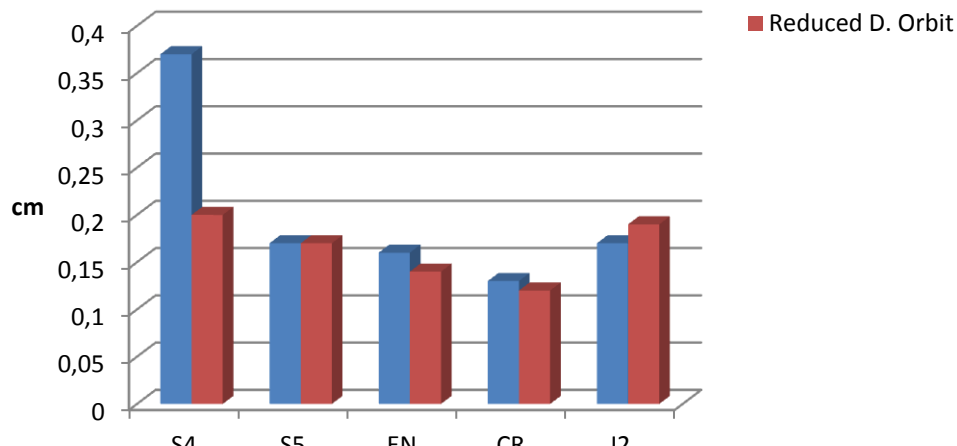
RMS radial



Mean radial



Std. dev. of Mean radial

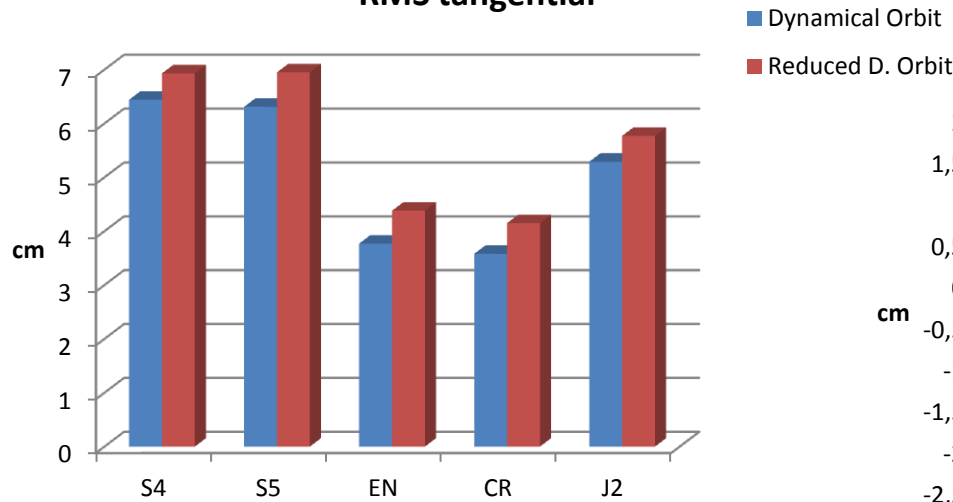


Comparison of the estimated DORIS orbits and CNES/SSALTO multitechnique

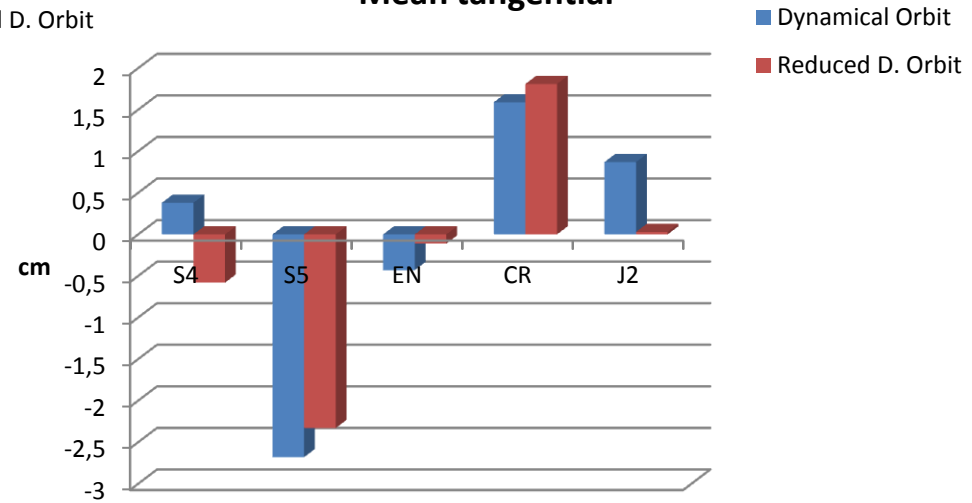
orbits - TANGENTIAL direction

Days 001-040 of 2011

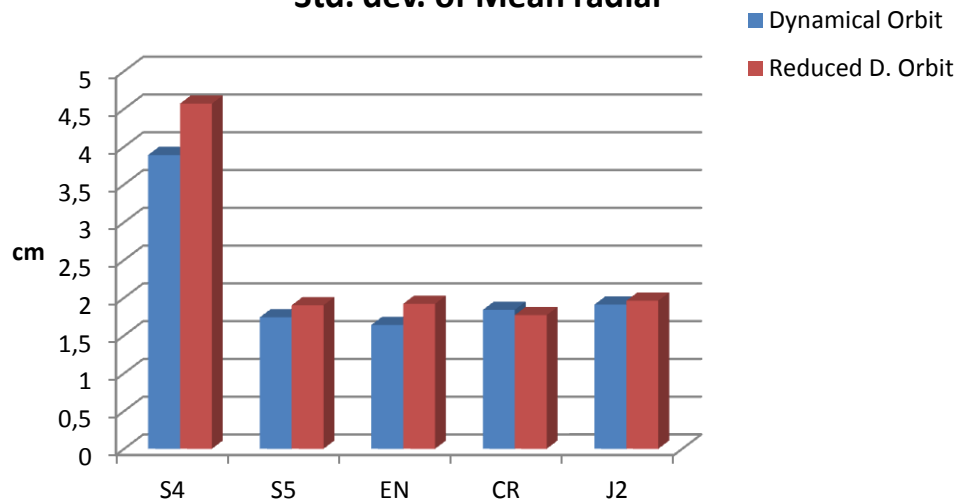
RMS tangential



Mean tangential



Std. dev. of Mean radial

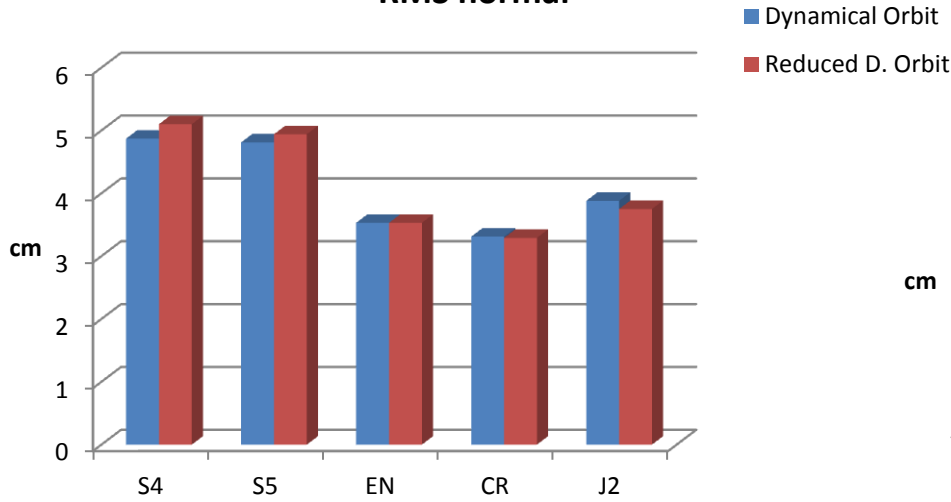


Comparison of the estimated DORIS orbits and CNES/SSALTO multitechnique

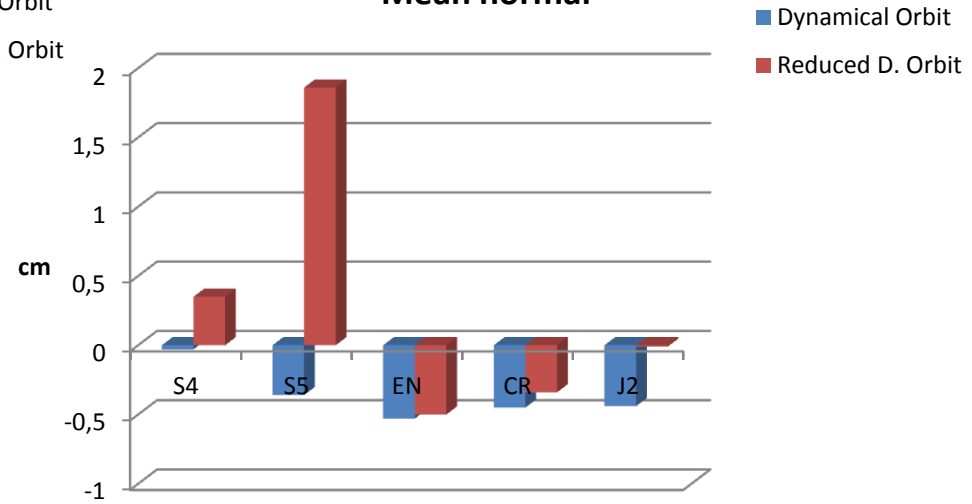
orbits - NORMAL direction

Days 001-040 of 2011

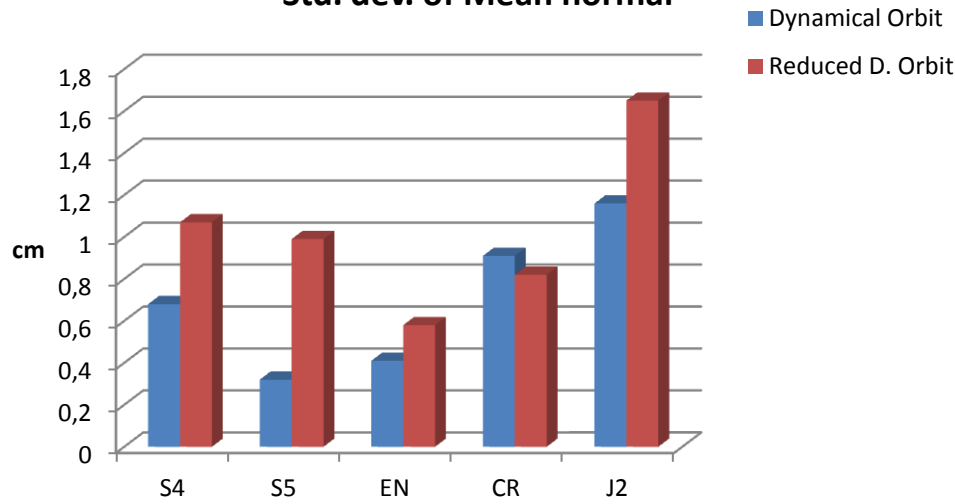
RMS normal



Mean normal



Std. dev. of Mean normal



RMS reduction of the orbit using dynamical modeling in comparison to reduced-dynamical. RMS with respect to SSALTO/CNES orbits as a reference

green – RMS decrement, red- RMS increment

	Radial	Tangential	Normal
SPOT-4	18%	7%	6%
SPOT-5	16%	6%	9%
Envisat	24%	14%	0%
Cryosat	28%	12%	1%
Jason-2	13%	7%	0%
<i>average</i>	20%	9%	3%

RMS total

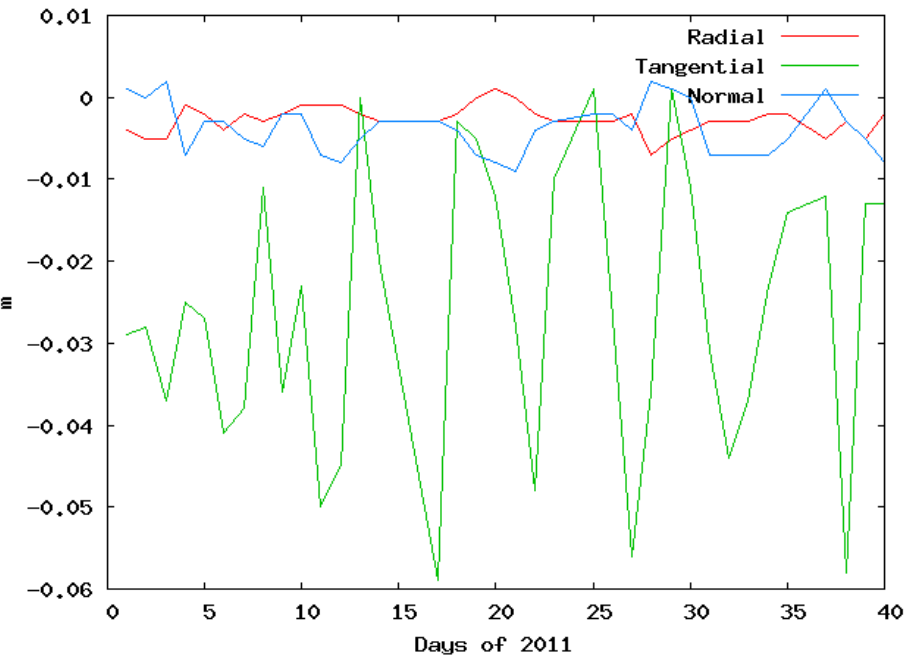
	Radial	Tangential	Normal
SPOT-4	22%	7%	5%
SPOT-5	17%	9%	3%
Envisat	24%	14%	0%
Cryosat	26%	14%	1%
Jason-2	15%	8%	3%
<i>average</i>	21%	10%	1%

RMS after Mean removal

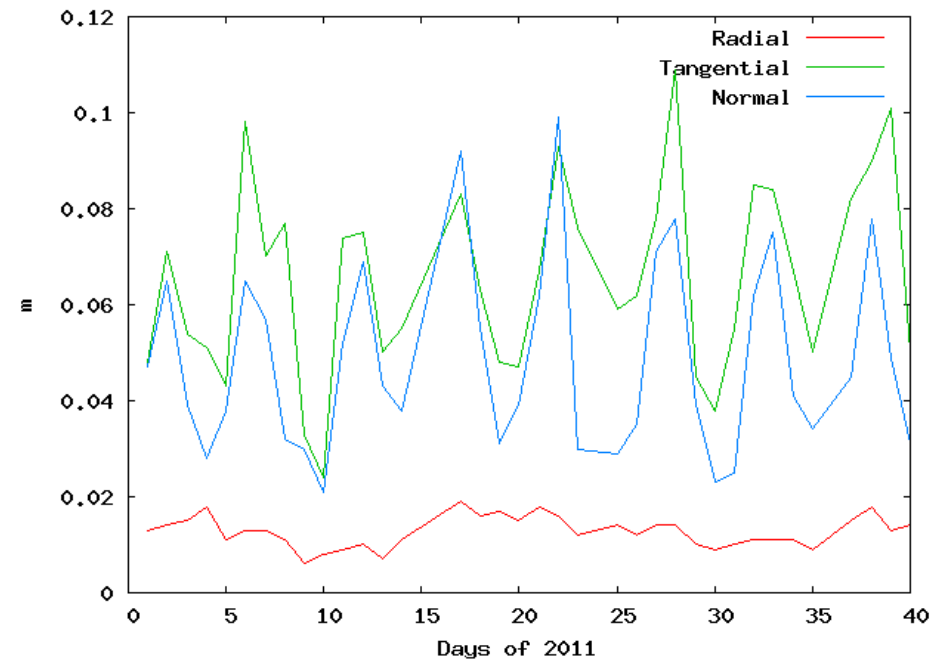
Comparison of the estimated DORIS orbits and CNES/SSALTO multitechnique orbits

- Days 001-040 of 2011
- Signal with period about 5 days strongly significant for SPOT-5, slightly for SPOT-4.
- Figure –daily Mean and RMS for SPOT-5

Mean Spot-5



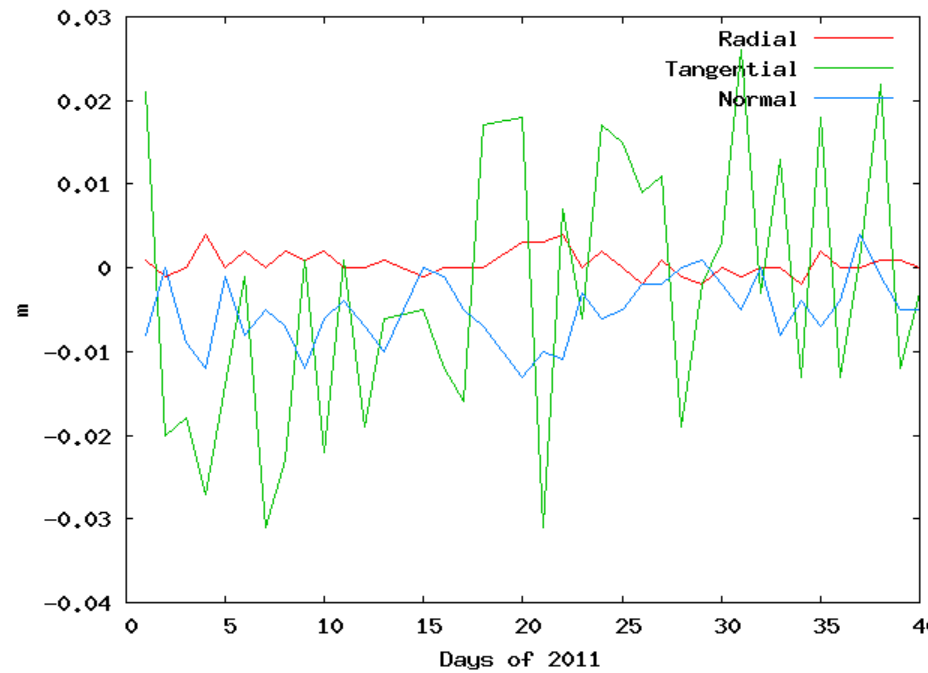
RMS Spot-5



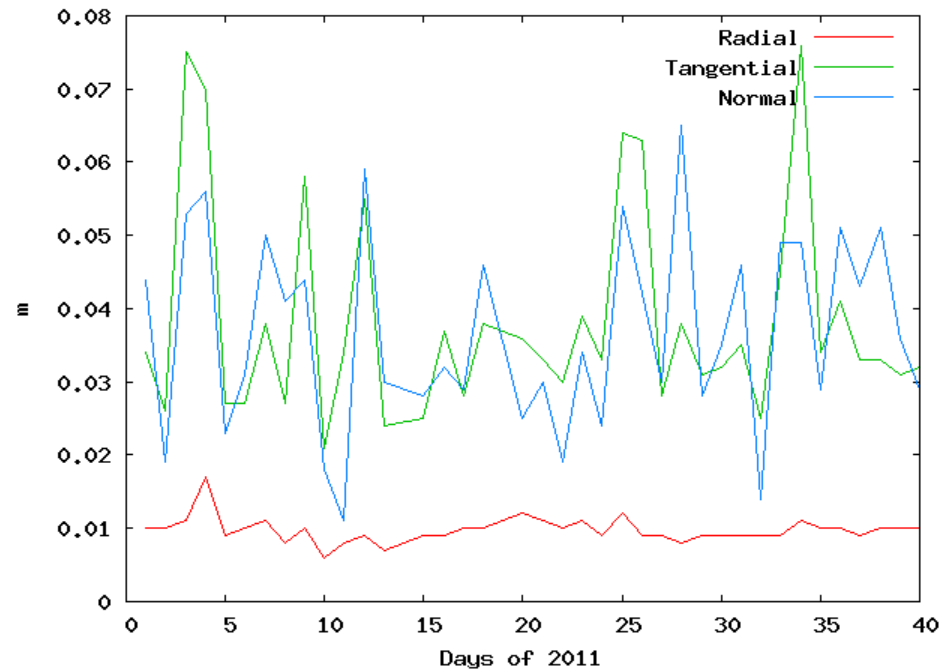
Comparison of the estimated DORIS orbits and CNES/SSALTO multitechnique orbits

- Figure –daily Mean and RMS for Envisat

Mean Envisat



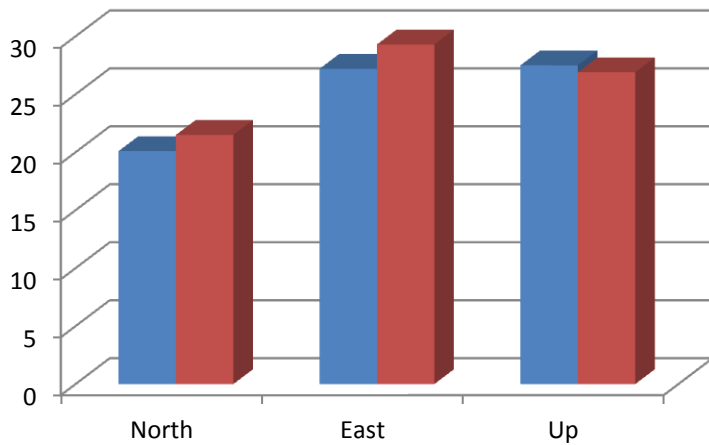
RMS Envisat



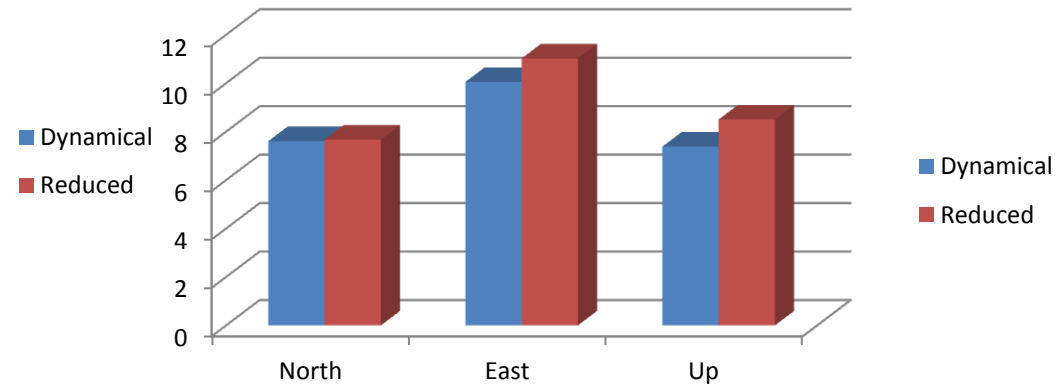
Free network solutions

- Weekly multi-satellite solutions
- third quarter of 2011

Comparison to DPOD08 (mm)



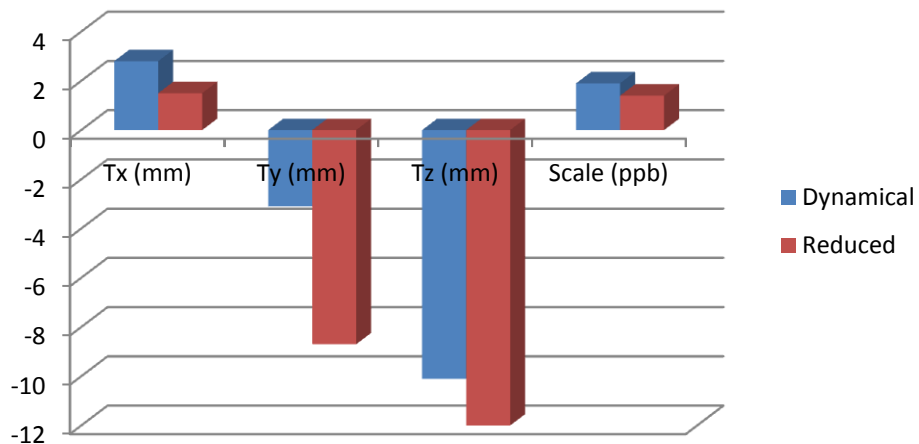
WRMS (mm)



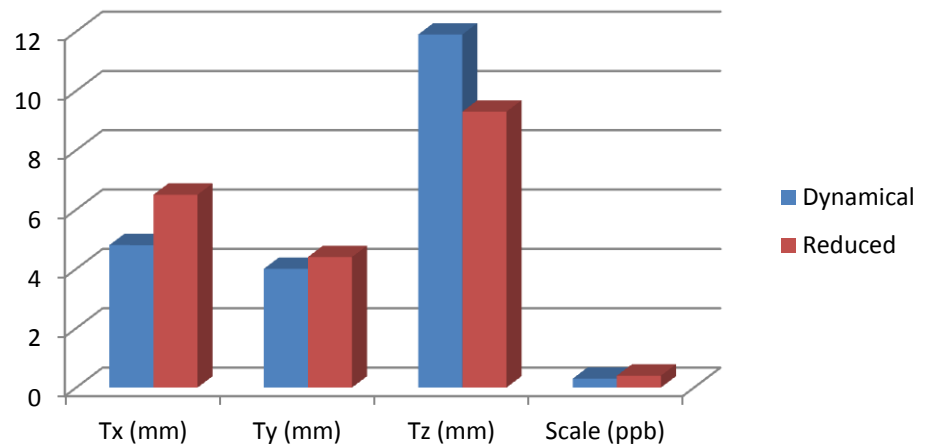
Free network solutions, TRF parameters vs. DPOD08

- Weekly multi-satellite solutions
- third quarter of 2011

Mean value of TRF par.



Variations of TRF par.



Earth Radiation Model

- Computation of Irradiance [W/m²] at satellite position, assuming:
 - Earth scattering properties approximated as a Lambertian sphere
 - Earth reflected radiation in the visible (albedo)
 - Earth emitted radiation in the infrared
- Types of models:
 - Analytical: Constant albedo, Earth as point source → large satellite altitude

$$\vec{E}_{ERM-A}(\psi, h) = \frac{A_E E_{sun}}{(R_E + h)^2} \left[\frac{2\alpha}{3\pi^2} ((\pi - \psi)\cos\psi + \sin\psi) + \frac{(1-\alpha)}{4\pi} \right] \hat{r}$$

$$A_E = \pi R_E^2, \quad R_E = 6378 \text{ km}, \quad E_{SUN} = 1367 \text{ W/m}^2, \quad h = \text{satellite altitude}, \quad \alpha = \text{albedo} (\approx 0.3)$$

For LEO satellites:

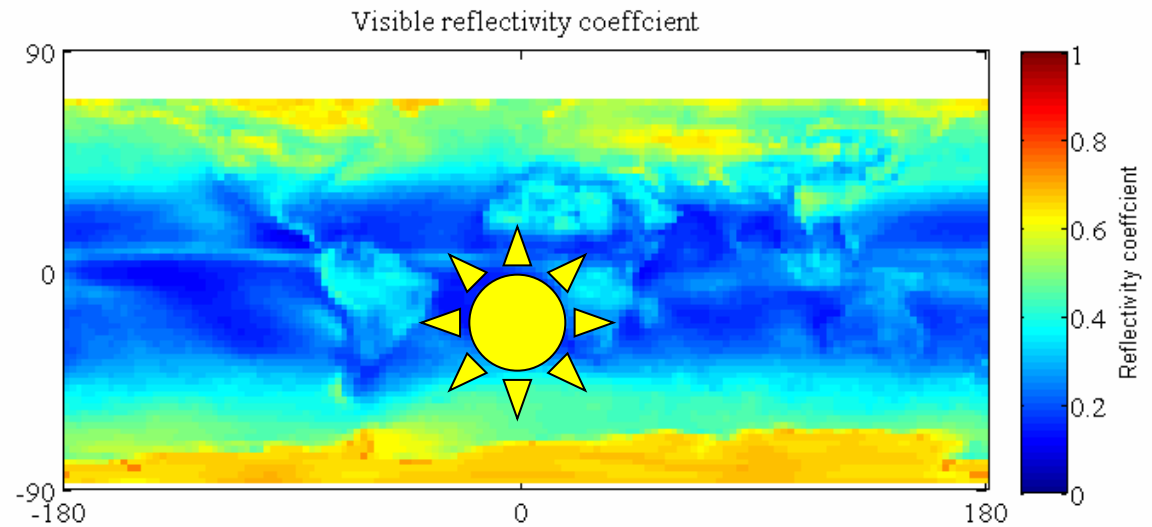
- Latitude- and time-dependent reflectivity and emissivity (Knocke et al., 1988)
- Latitude-, **longitude**- and time-dependent reflectivity and emissivity from NASA CERES project

Earth Radiation Model

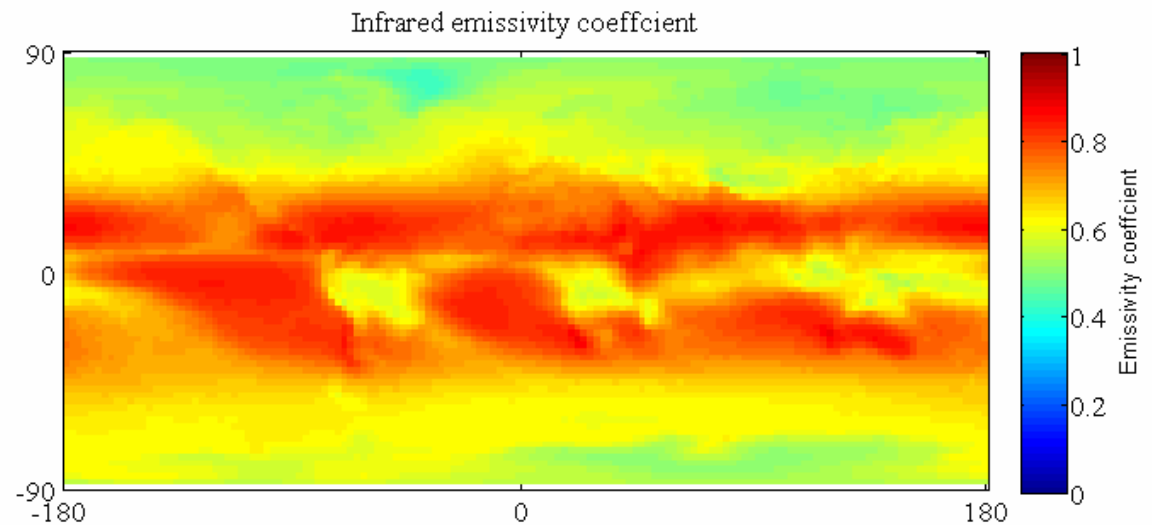
CERES

(Clouds and Earth's
Radiant Energy System)
NASA EOS project

Reflectivity (visible) →



Emissivity (infrared) →



CERES data, average (2000-
2010) for January

Earth Radiation Model

Irradiance [W/m^2]

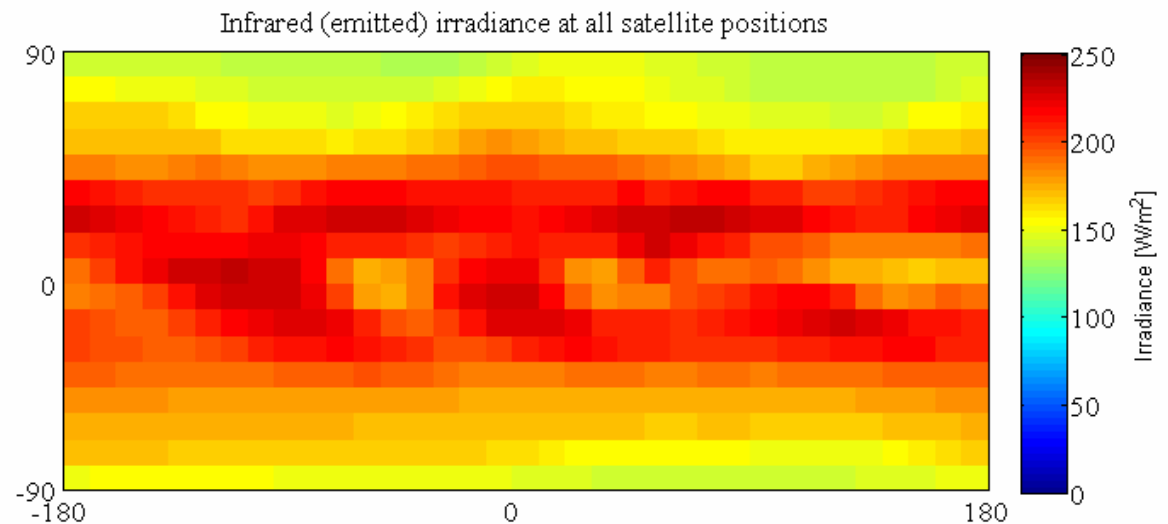
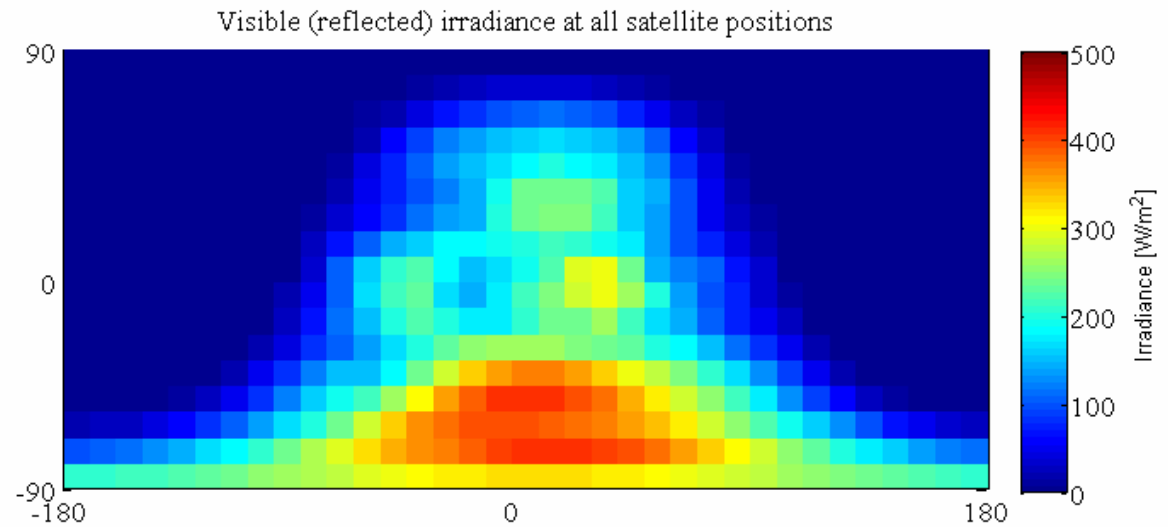
Average values for 12:00
UTC, January 2000-2010

received by a satellite
at 700 km altitude
at all possible positions

Reflected irradiance →
(reflected sunlight) in the
visible

Emitted irradiance → in
the infrared

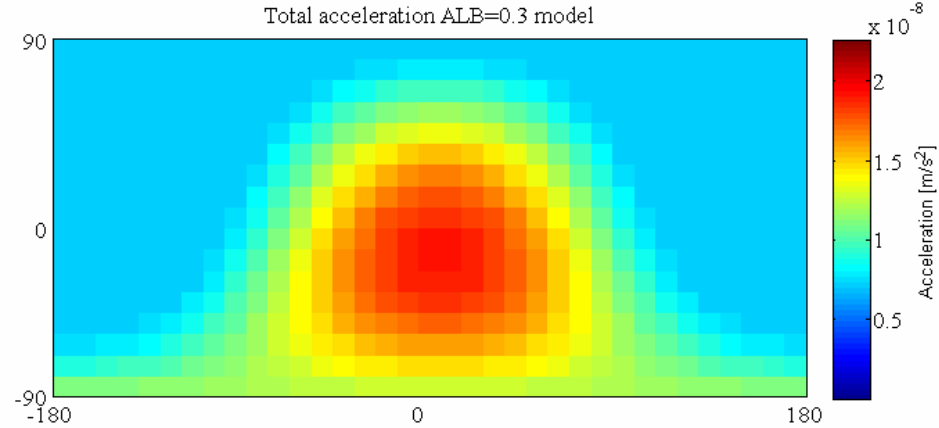
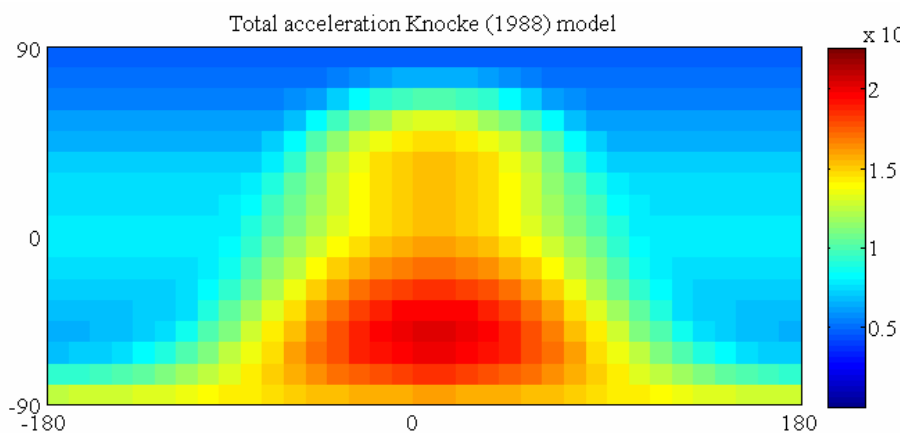
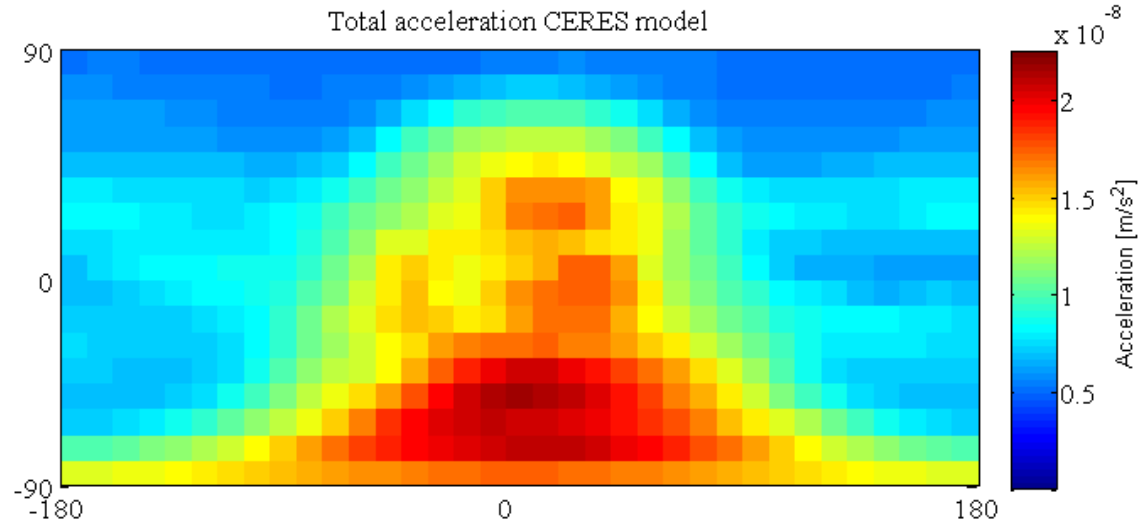
Integration of irradiance
over the area of the Earth
visible by the satellite



Acceleration due to Earth Radiation

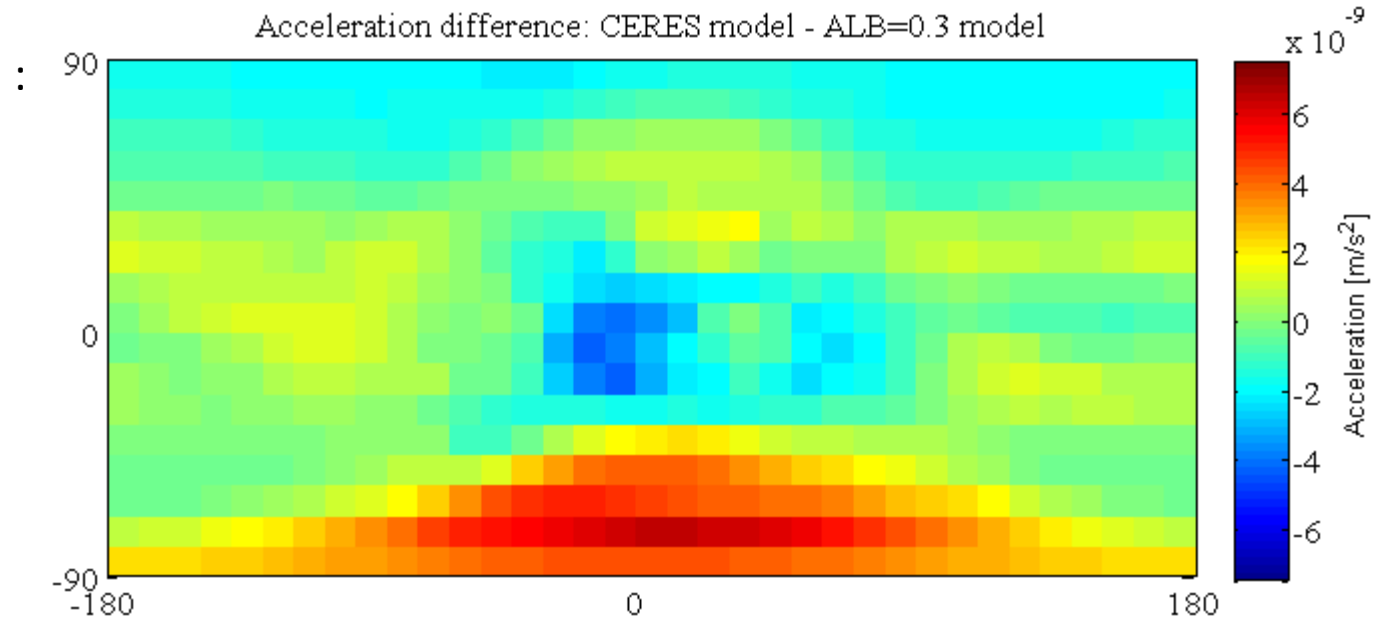
Radial acceleration using: area-to-mass ratio of $0.011 \text{ m}^2/\text{kg}$ (CRYOSAT-2)

Average values for 12:00 UTC, January 2000-2010

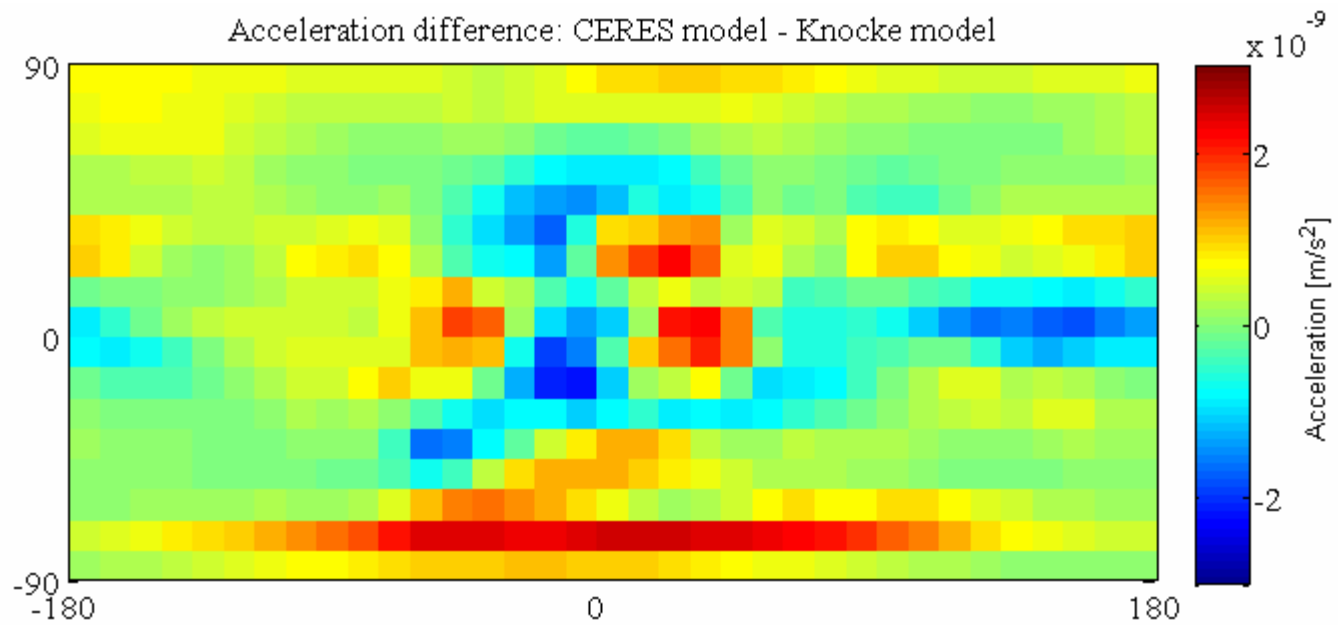


Acceleration due to Earth Radiation – model differences

Acceleration difference: CERES model - ALB=0.3 model



Acceleration difference: CERES model - Knocke model



FUTURE prospects

Dynamical orbit modeling

- complex testing, long time series
- comparison with external orbits (SSALTO and others)
- internal orbit overlaps
- SLR validation
- Impact on the free network solutions
- Results will be partially presented at IDS DORIS workshop (September 2012), complex presentation planned for AGU Fall meeting (December 2012)

Earth radation

- Studies of Analytical, Knocke and CERES models
- To be presented at IDS DORIS WORKSHOP (September 2012)